Residual Stellar Biases in Transmission Spectroscopy:

PLATO – ESP2025 Planets throughout the habitable zone

Marseille, Monday June 23, 2025

William Dethier

A Caution for Atmospheric Characterization of Planets in the Habitable Zone



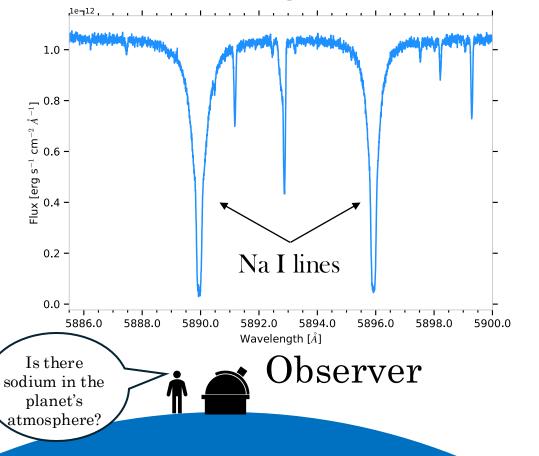


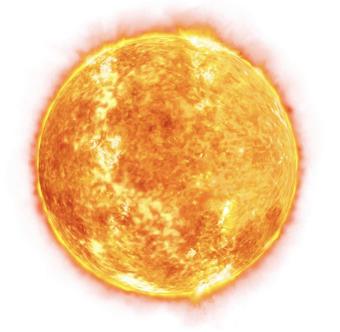






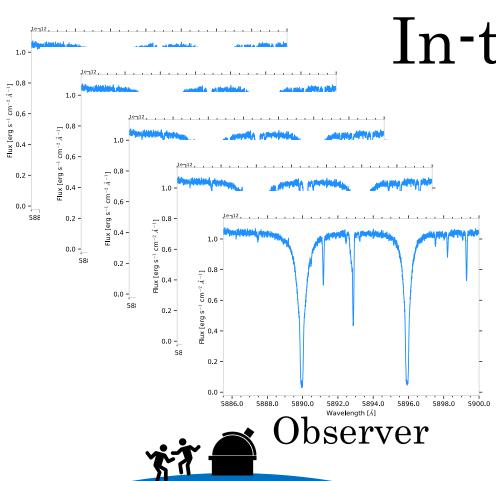
Out-of-transit spectrum



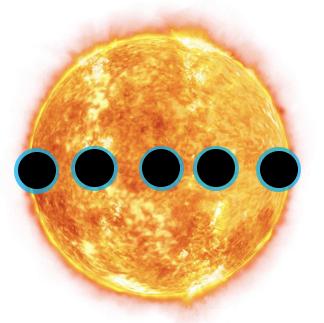


Target star

Observed before/after the transit



In-transit spectra



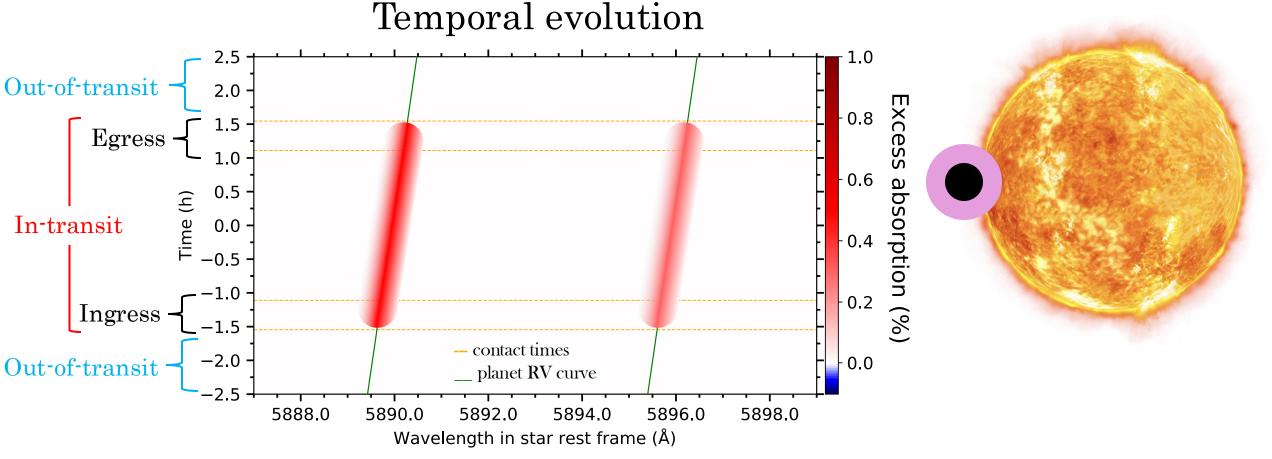
Target star

Observed during the transit

Absorption spectrum

$$A = \frac{F_{out} - F_{in}(t)}{F_{out}}$$

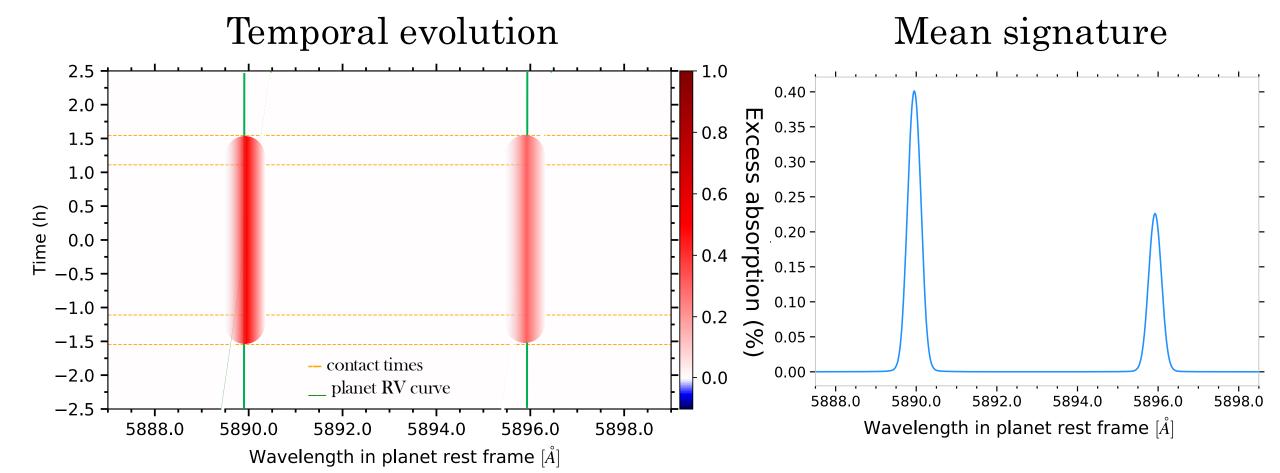
What we expect to see...



Absorption spectrum

$$A = \frac{F_{out} - F_{in}(t)}{F_{out}}$$

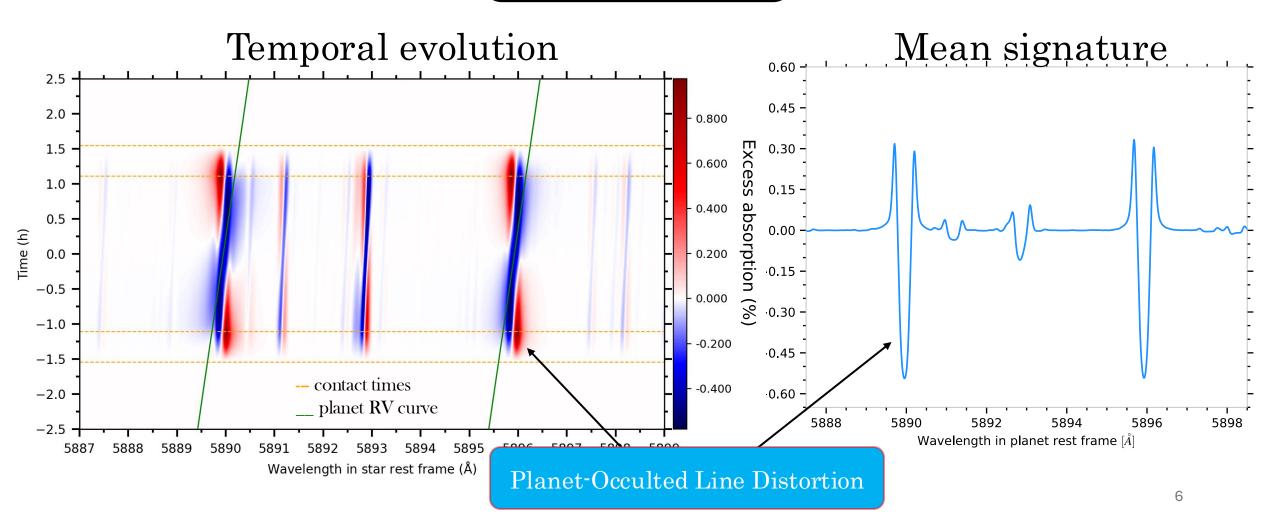
What we expect to see...



Absorption spectrum

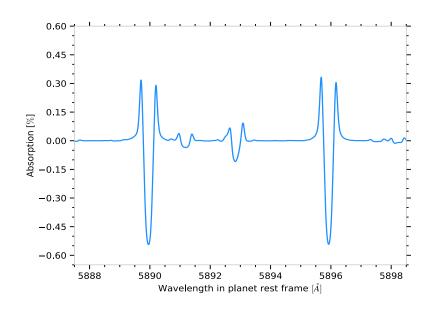
$$A = \frac{F_{out} - F_{in}(t)}{F_{out}}$$

What we actually see



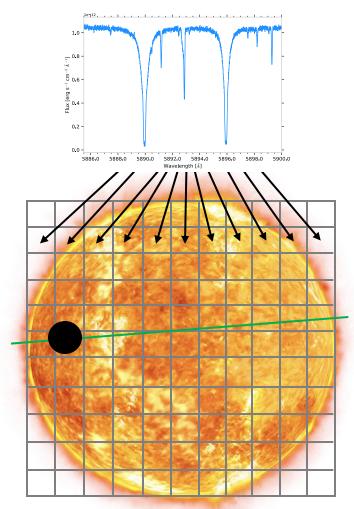
How to deal with these?

- 1. Distribute local spectra on grid
- 2. Simulate a transiting bare planet
- 3. Compute synthetic absorption spectrum

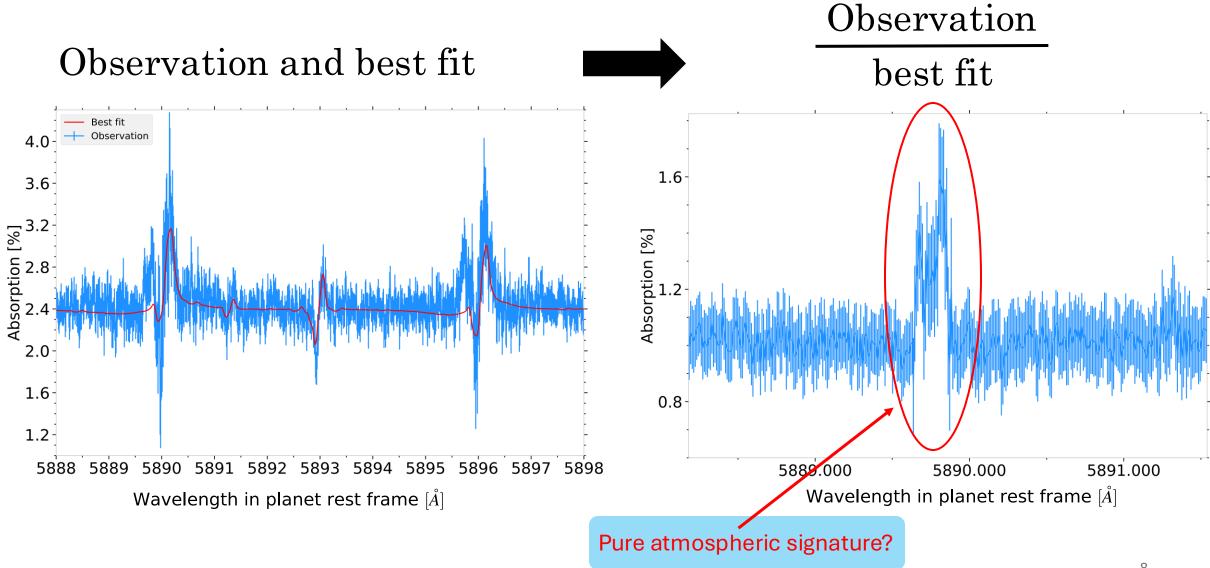


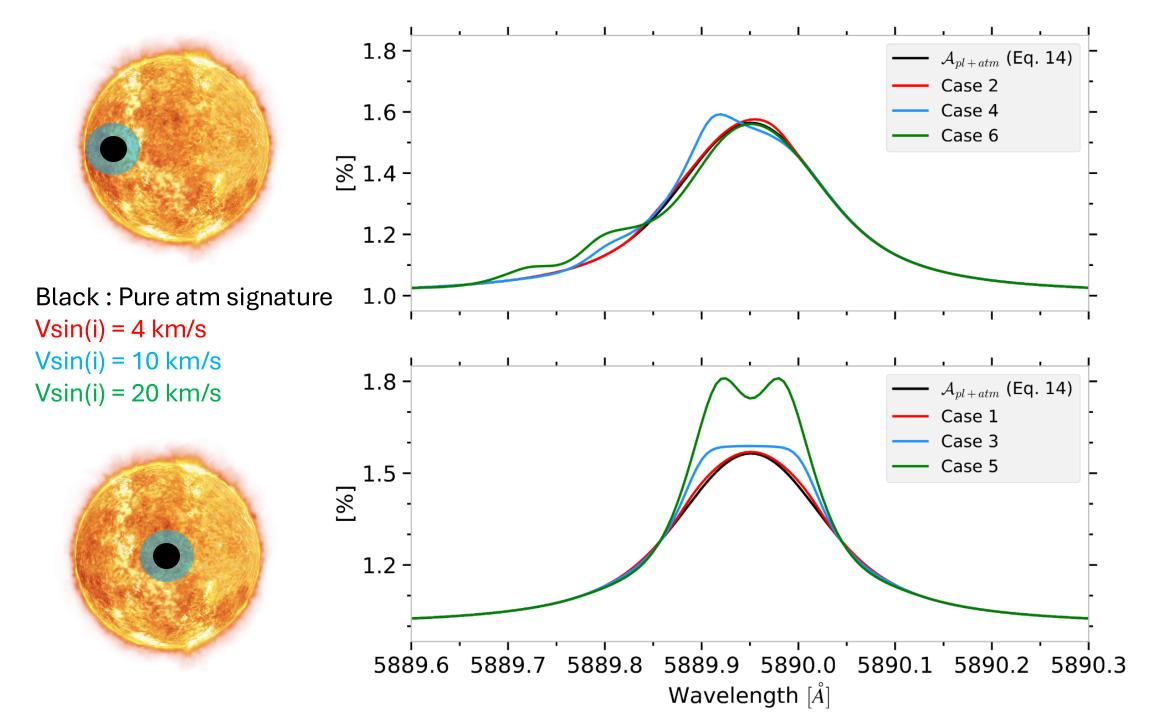


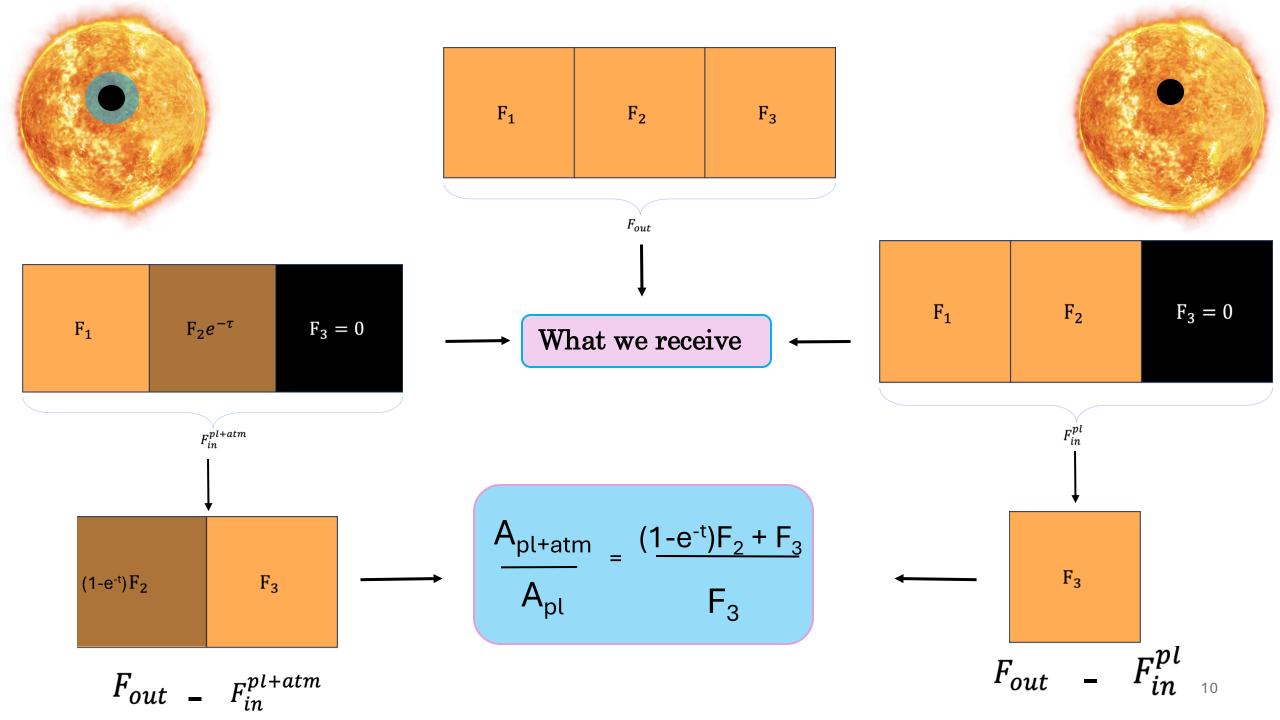
- 4. Use simulated spectrum to correct observations
 - Analyse residual atmospheric signal



Correcting absorption spectra

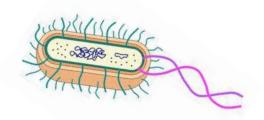




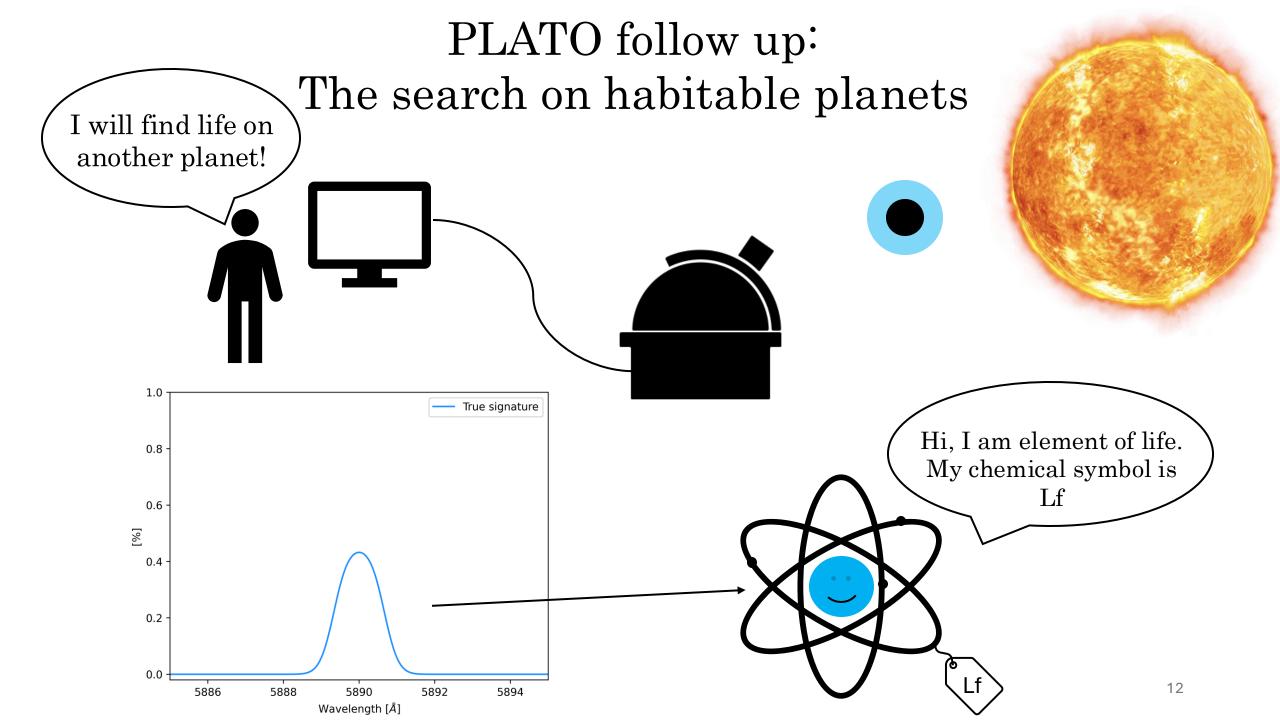




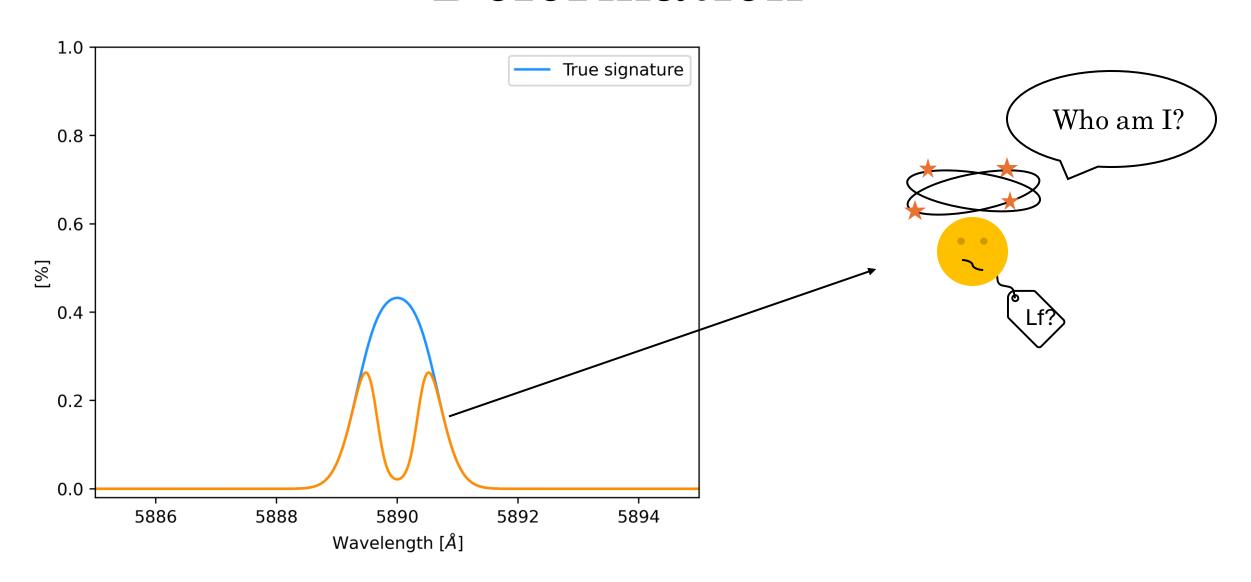
Fine, but what is the link with PLATO?



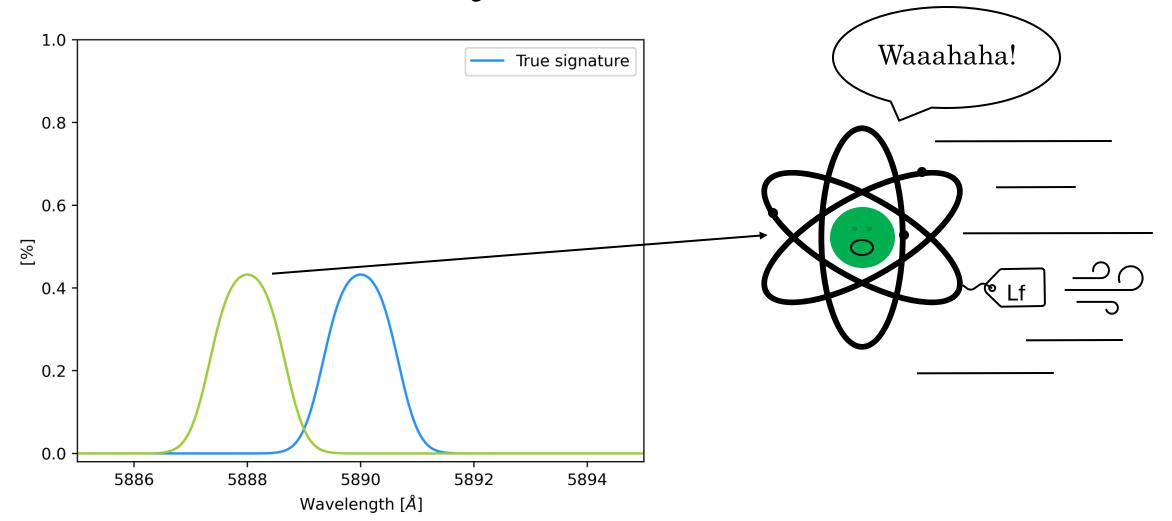




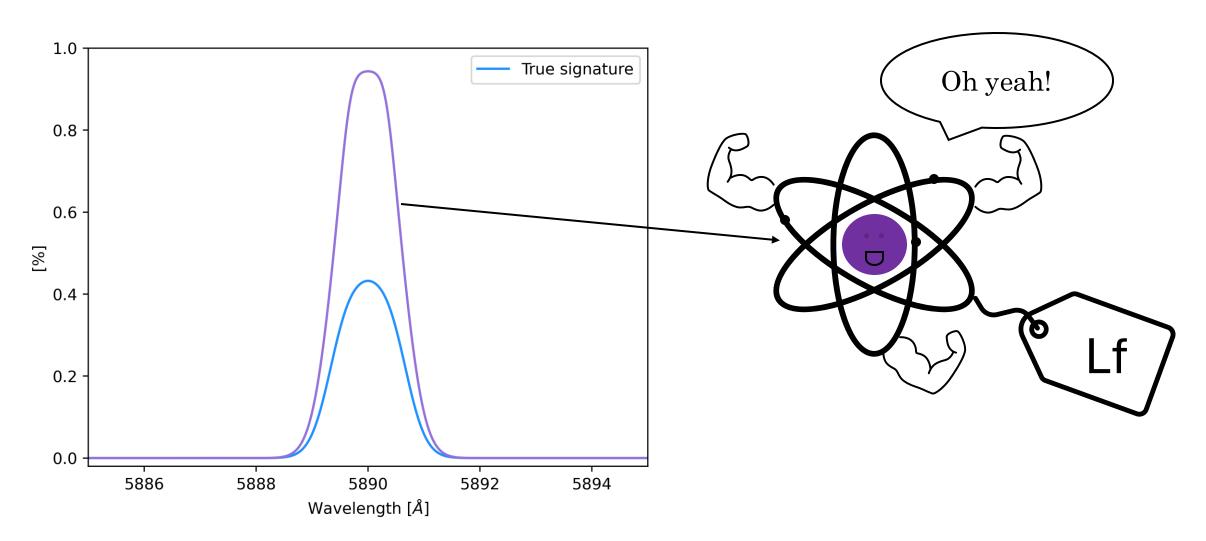
Deformation



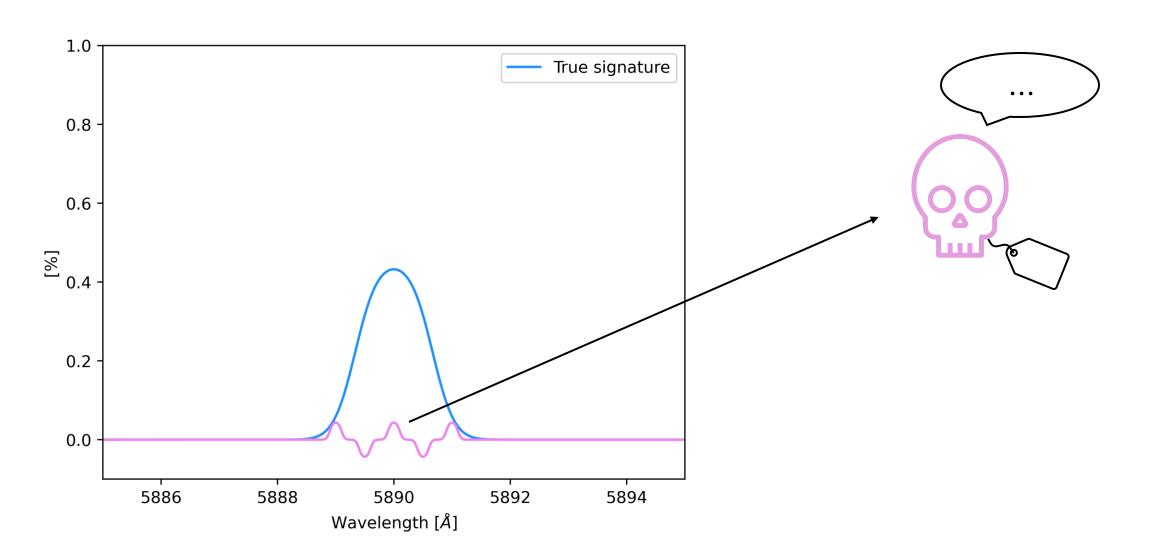
Dynamics



Overestimation

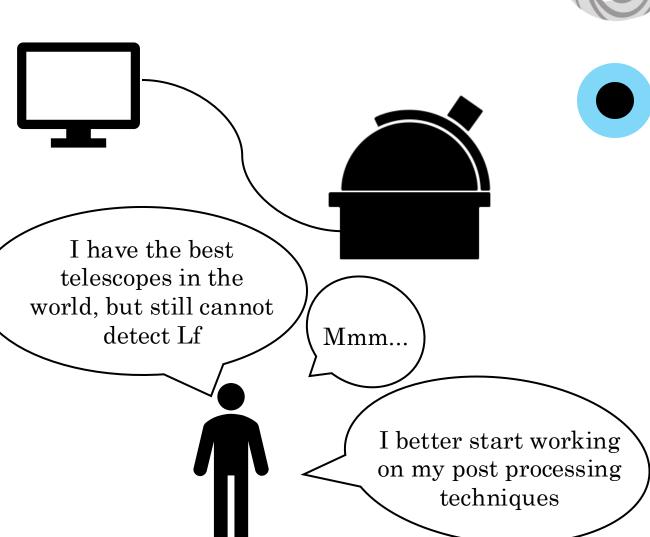


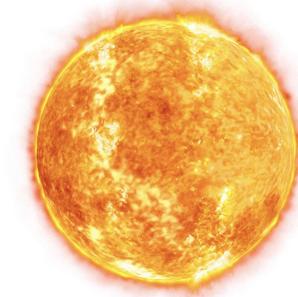
False non detection



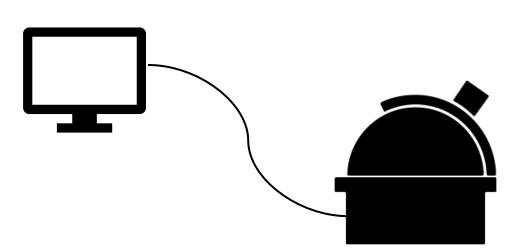
Conclusions



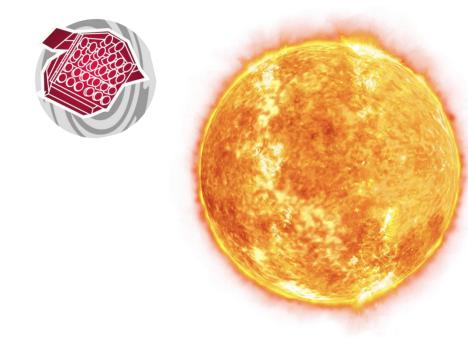




Conclusions



leanwhile on Lf0001b...



Did you forget the address on the invitation?
The Earthians cannot find us



