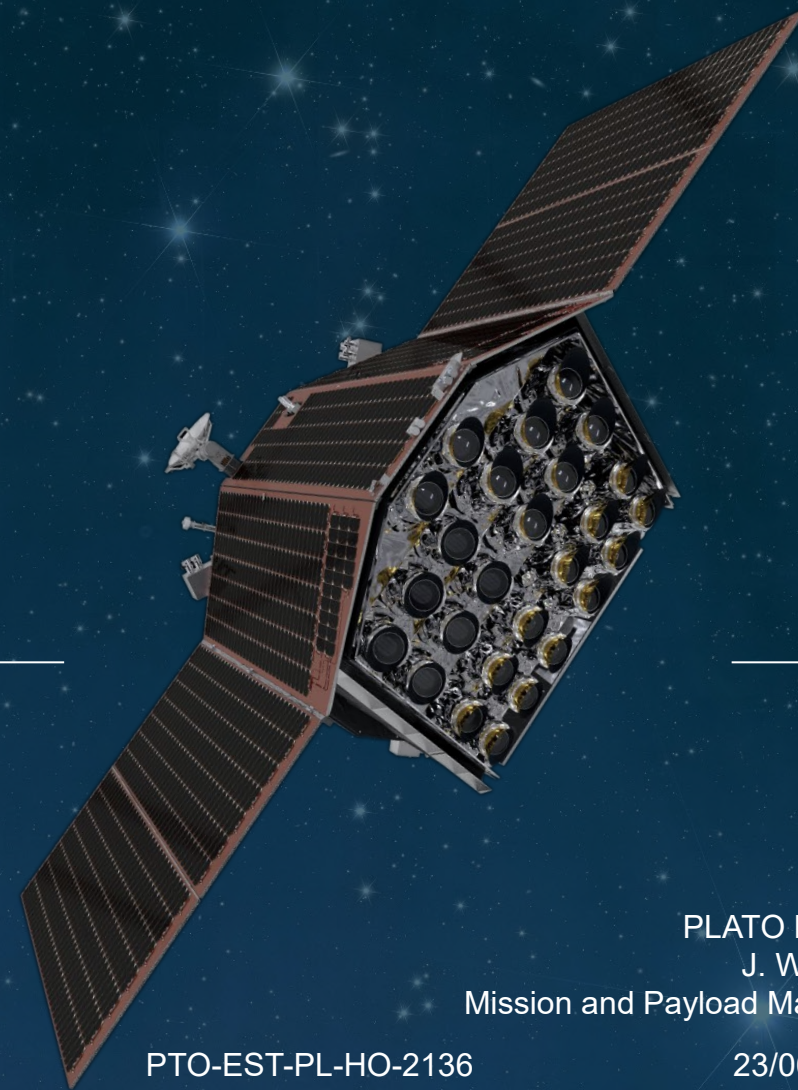




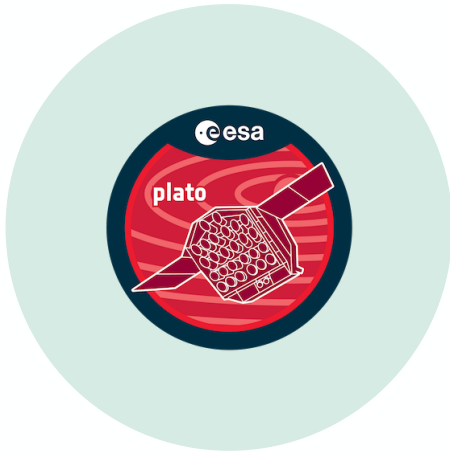
PLATO Instrument Status ESP2025



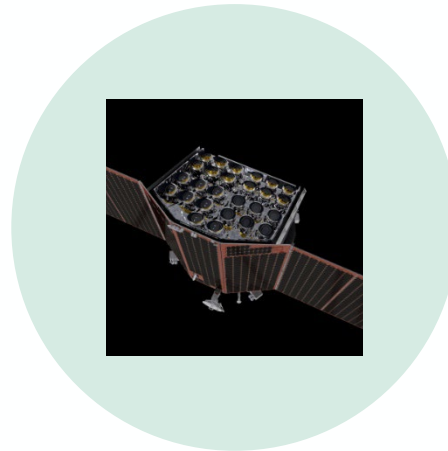
PLATO Project
J. Windsor
Mission and Payload Manager

PTO-EST-PL-HO-2136

23/06/2025



MISSION OVERVIEW



SPACECRAFT STATUS



PAYLOAD STATUS



Mission Overview

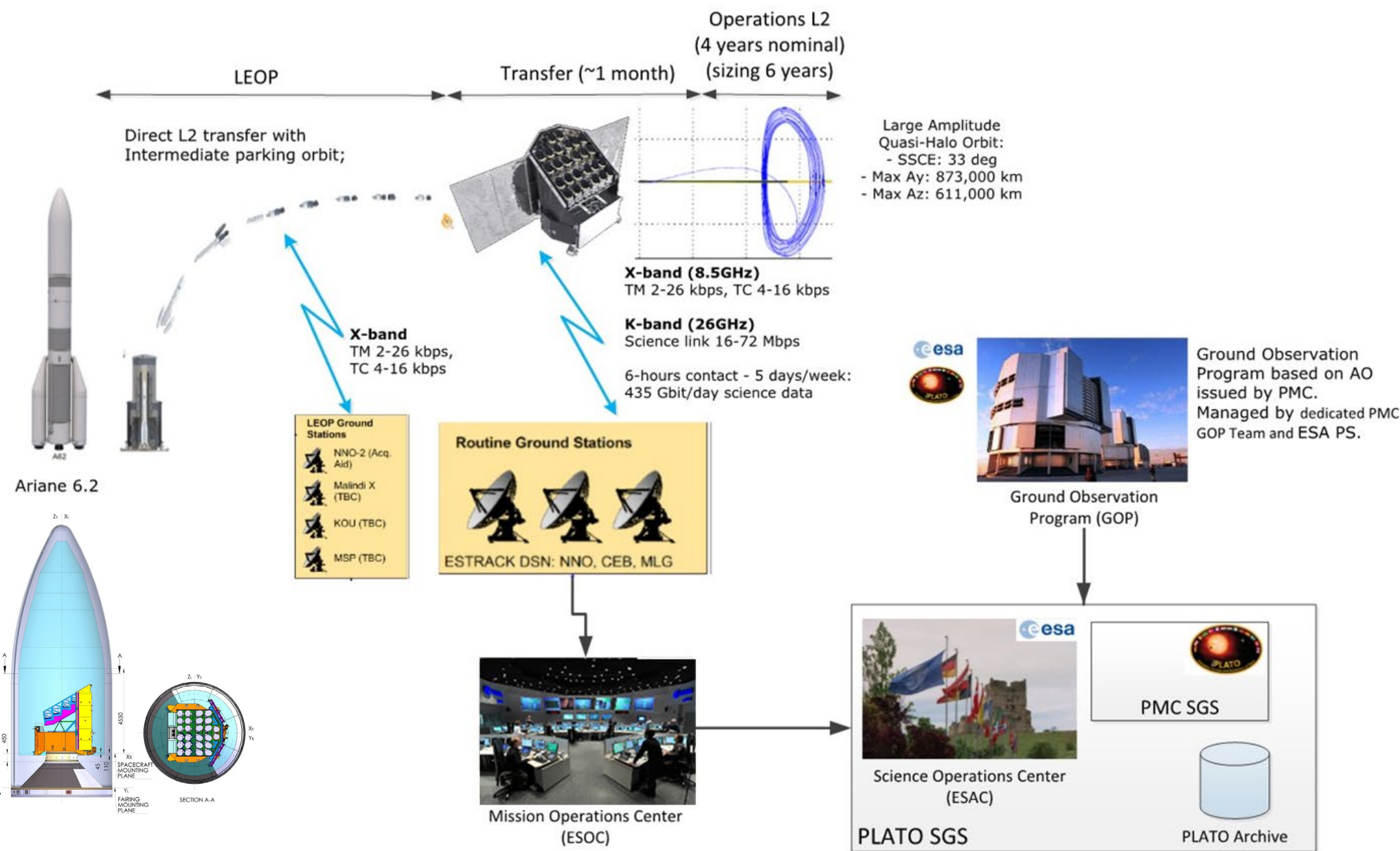
PLATO Mission – M3

PLAnetary Transits and Oscillations of stars (PLATO) is a mission to detect and characterise exoplanets and study their host stars

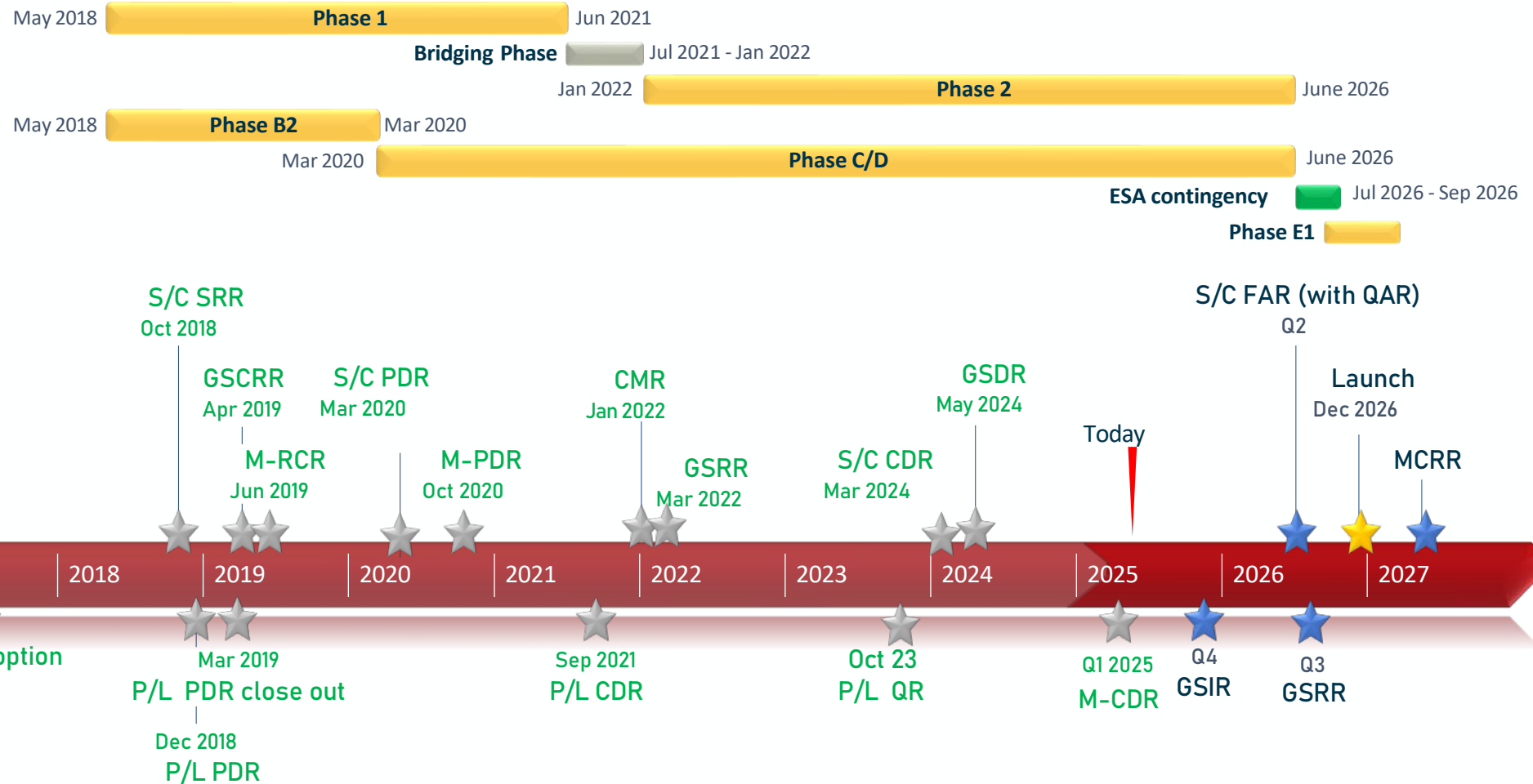
Focus on Earth-size planets in orbits up to the habitable zone of bright Sun-like stars to address 3 main questions:

1. How do planets and planetary systems form and evolve?
2. Is our Solar system special or are there other systems like ours?
3. Are there potentially habitable planets?

PLATO Mission Architecture



PLATO Milestones Status

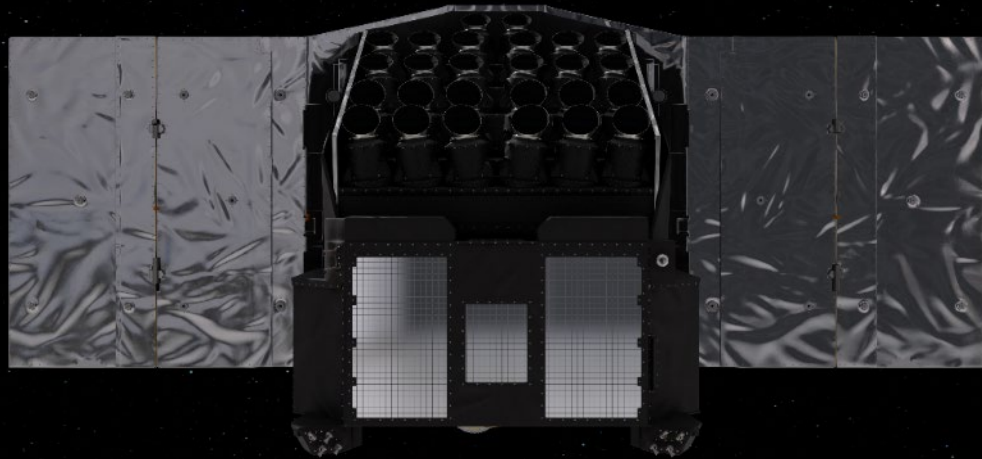
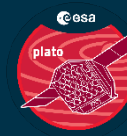




Spacecraft Status



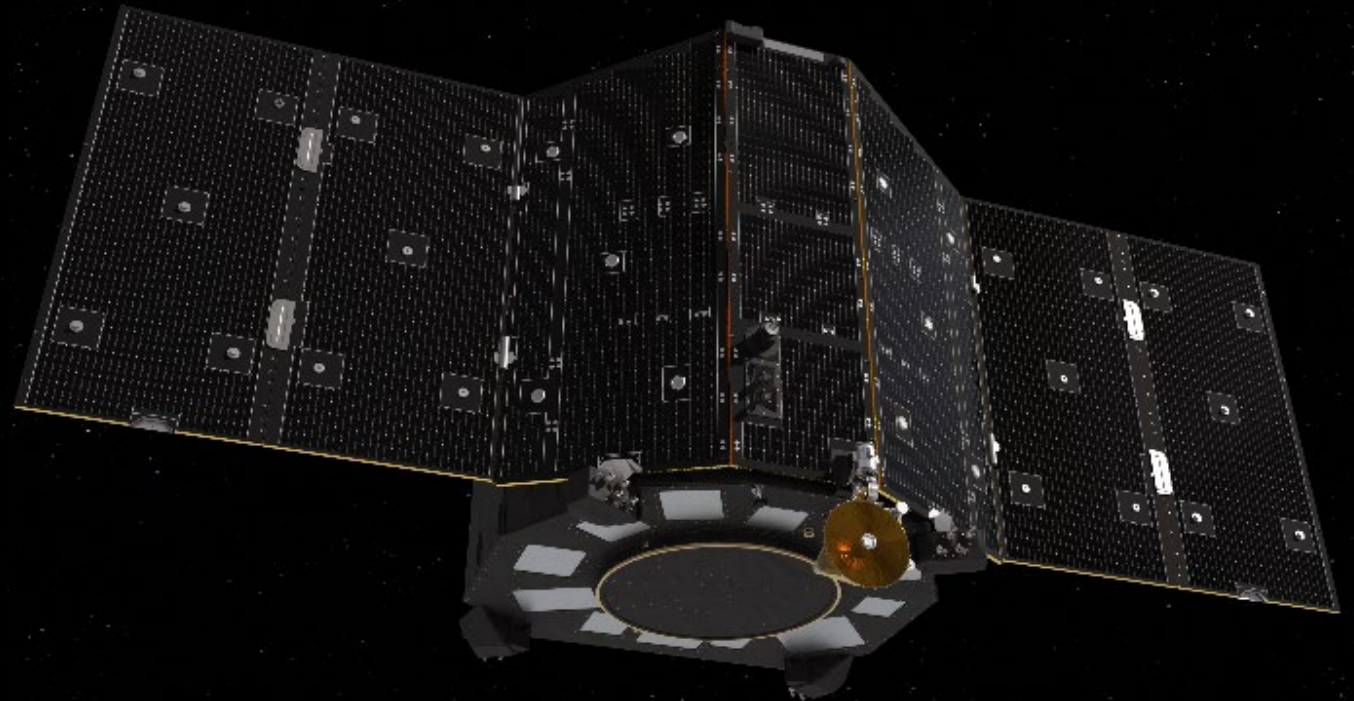
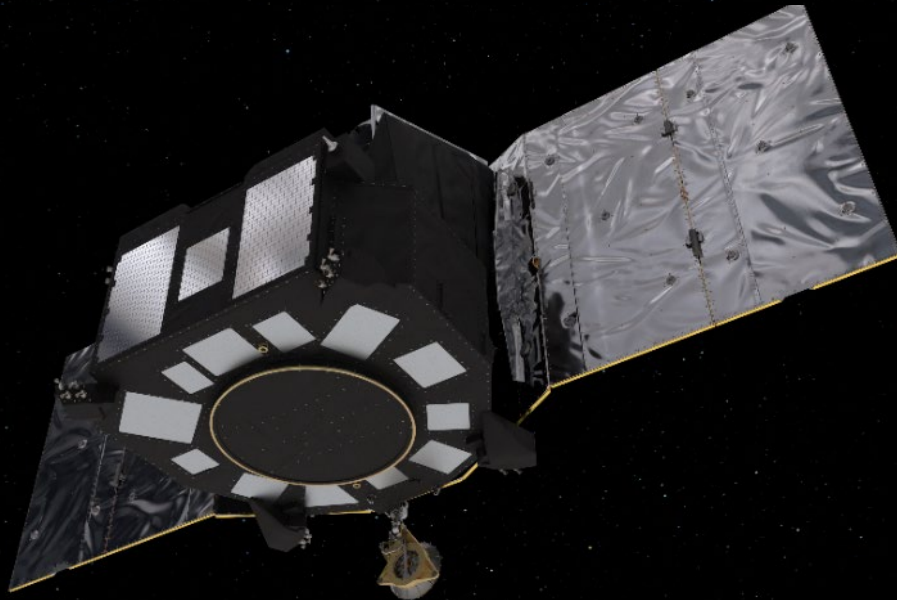
PLATO Space Segment



Spacecraft mass and power generation:

2213* kg Spacecraft at launch (w/o system margin)
2595 kg Launcher Performance (SRD and SC sizing mass)
2799 W min available power from SA for nom science ops
vs 2629W max power required (incl. 15% sys margin)

*including 80 kg LVA and 172.5 kg propellant and 10 Kg balancing mass



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PLATO Camera Alignment

Y [arcmin]

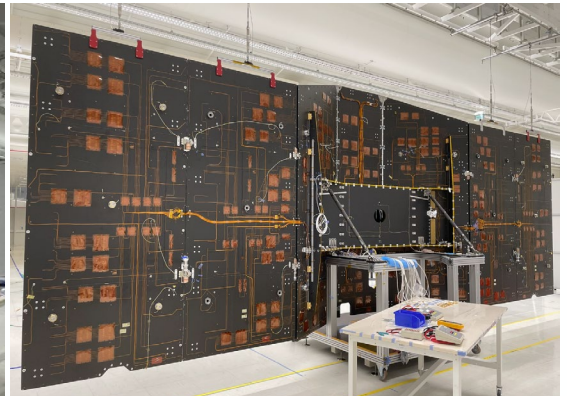
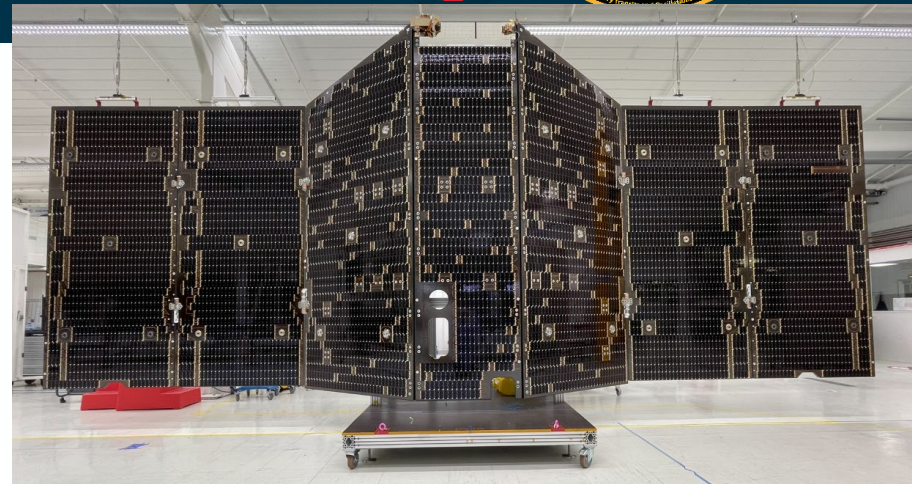
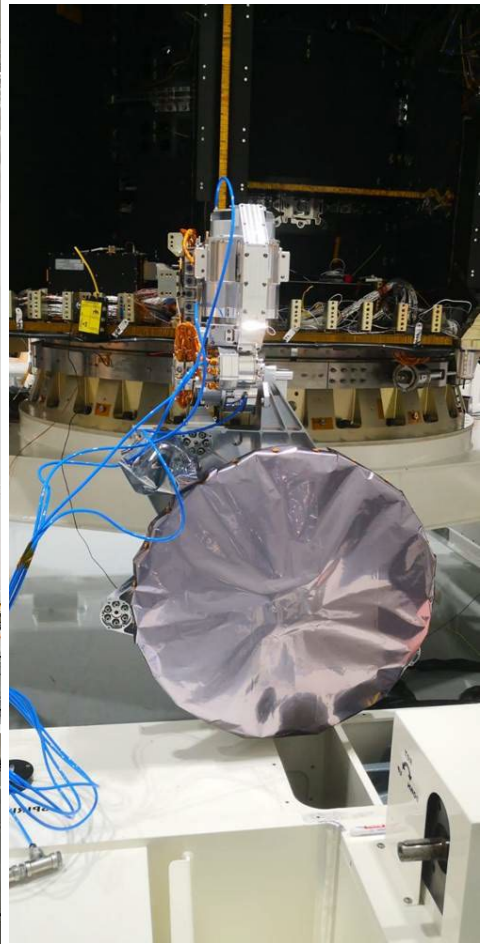
X [arcmin]

- Group 1
- Group 2
- Group 3
- Group 4
- F-CAM Blue (FM19)
- F-CAM Red (FM23)
- Requirement
- Mean of N-CAMS

PLATO – Spacecraft PFM Hardware



Service Module now at OHB

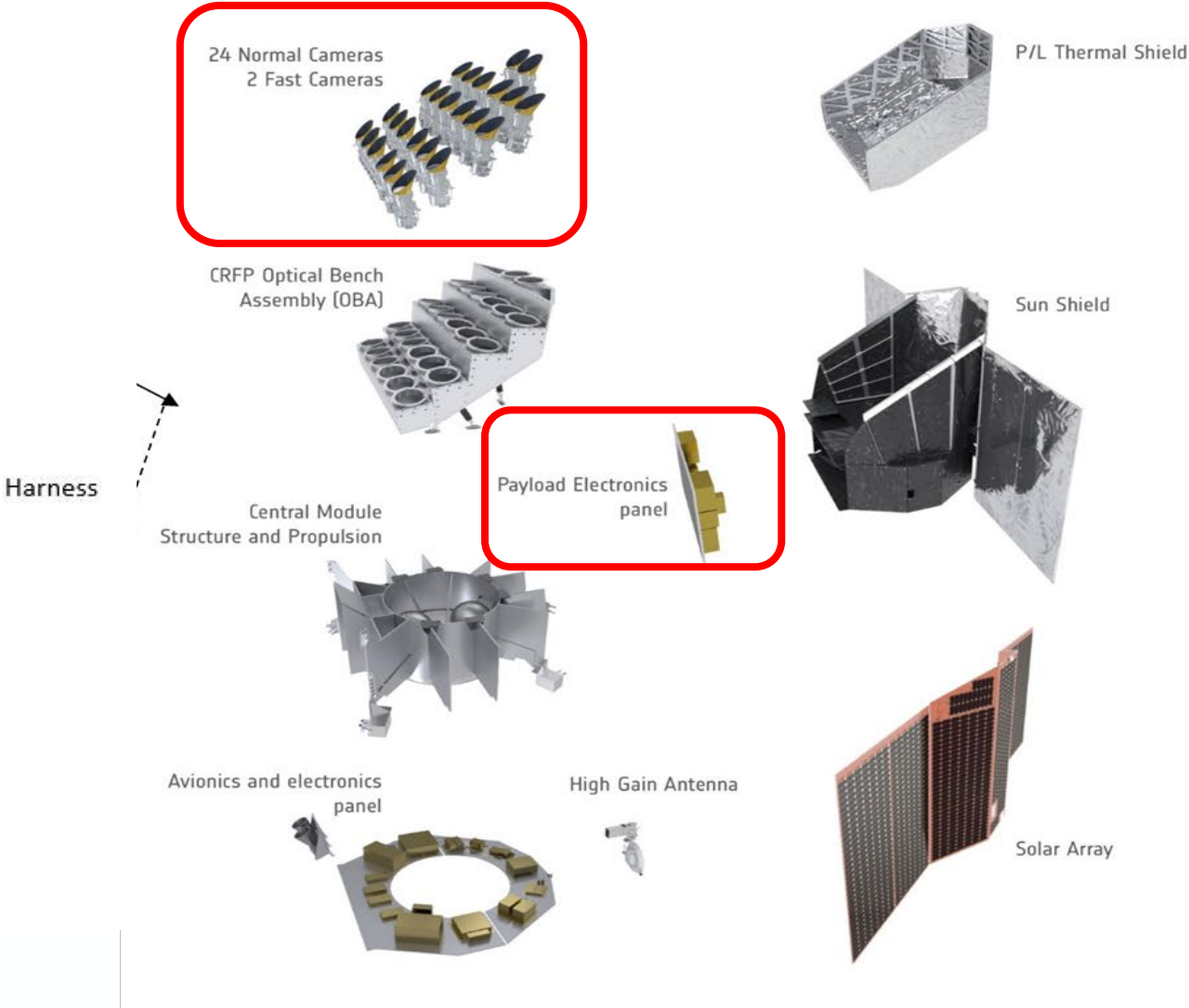
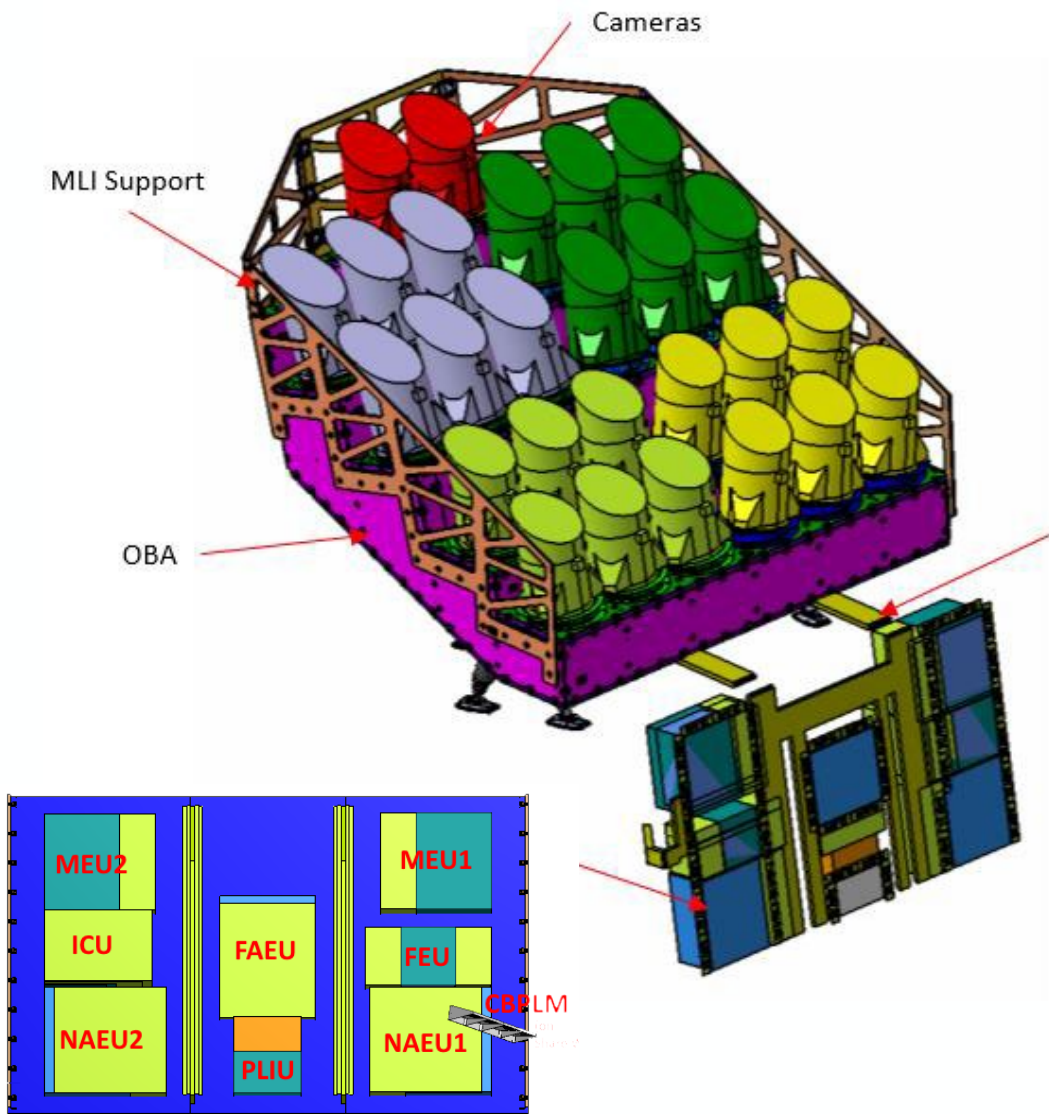


Sunshield/Solar Array

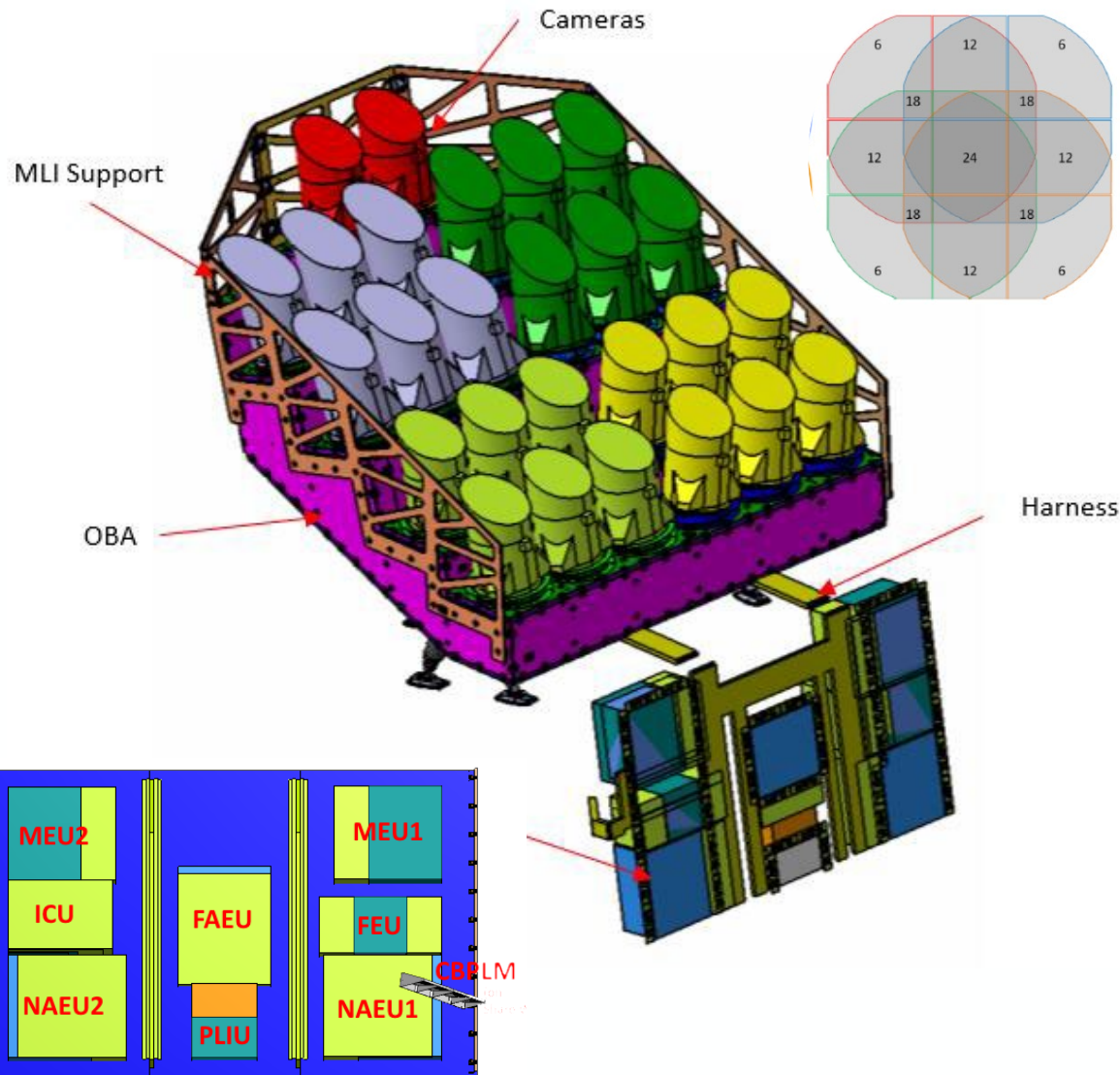


Payload Status

Payload Module Overview (1/2)



Payload Module Overview (2/2)



Instrument concept

- Wide Field-of-View and large photometric dynamic range to maximise the number of observable stars
- Multi-telescope configuration with CCD-based focal planes in the visible wavelength

CAMERA Subsystem (CAM)

- 24 almost identical “normal” cameras (N-CAM)
- 2 “fast” cameras (F-CAM) for observation of brighter stars and Fine Guidance Sensor capabilities (AOCS)
- Ancillary Electronic Units (2x N-AEU & 1x F-AEU)

Data Processing Subsystem (DPS)

- Main Electronic Units (2x MEU) for N-CAMs
- Fast Electronic Unit (FEU) for F-CAMs
- Instrument Control Unit (ICU)

PLATO Payload Overview

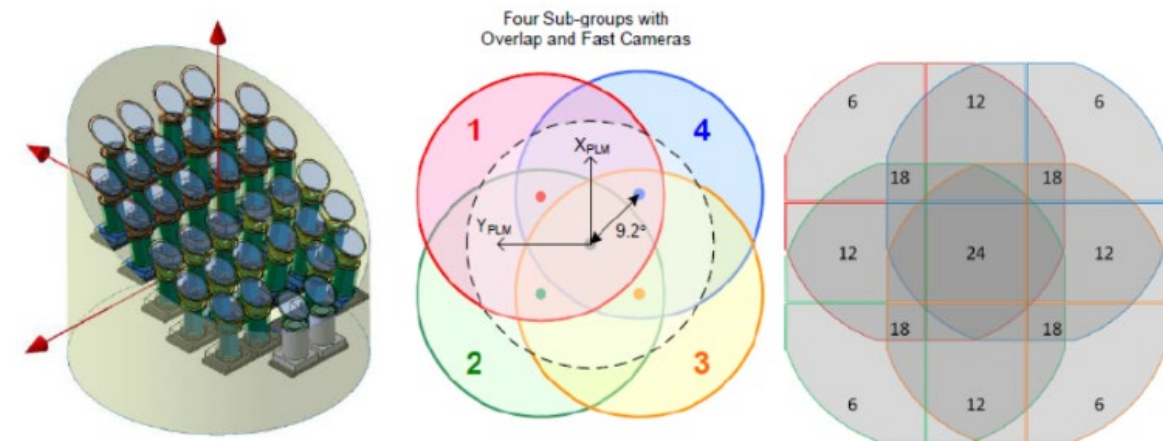


Driving payload requirements

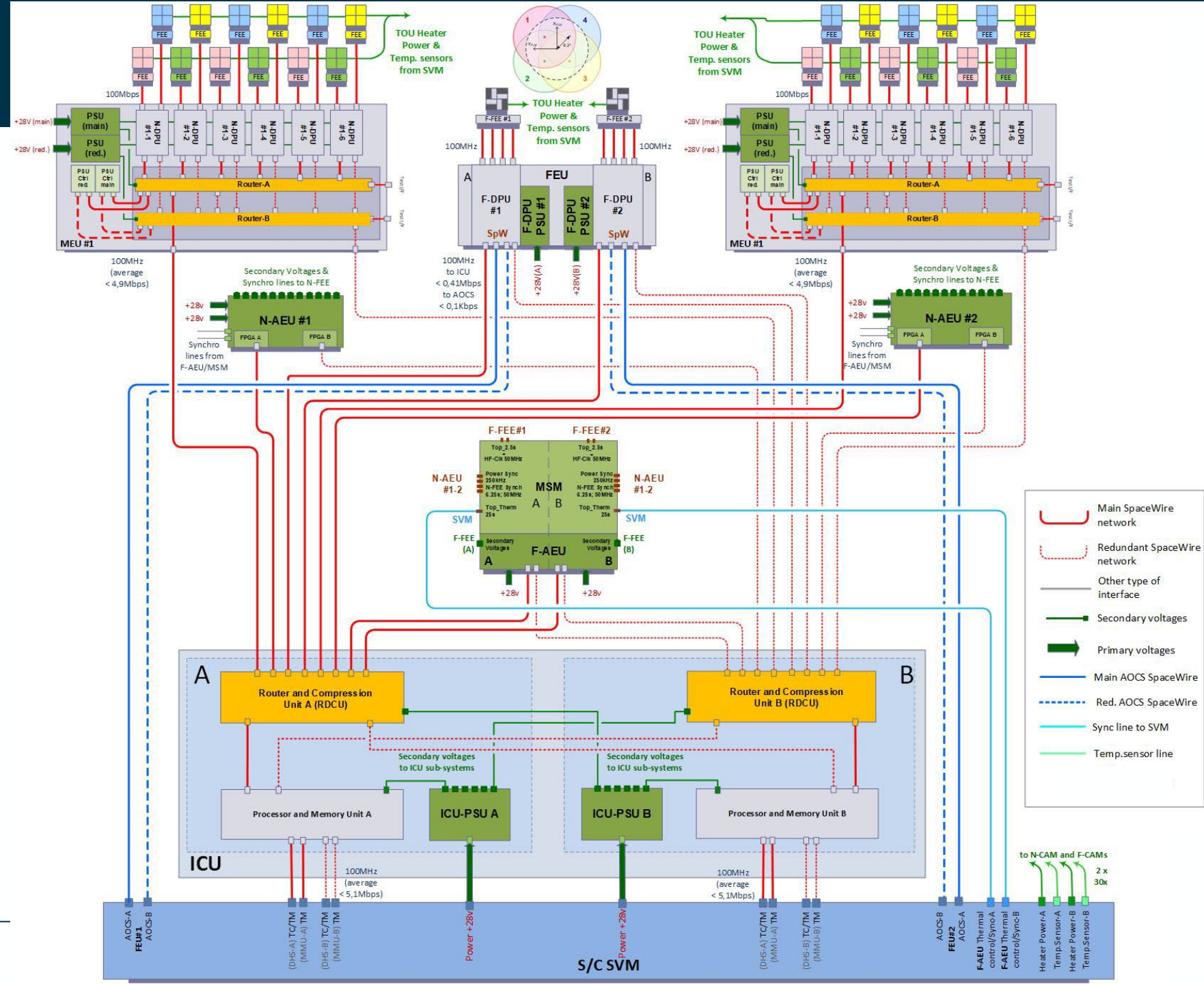
- ❑ Optimization of Field-of-view to maximize number of bright stars observed simultaneously
- ❑ Photometric accuracy (signal-to-noise and resolution) needed to detect Earth-sized planets and photometric dynamic range (bright stars with apparent magnitude $m_v < 11$ as well as fainter stars down to $m_v = 16$)
- ❑ Attitude performance, stability and knowledge to reduce error in the measured light-curves
- ❑ Level of on-board processing to limit data volume
- ❑ Ground segment processing requirements and algorithms for data reduction
 - 4 Groups of 6 normal cameras
 - Each group with same FoV and 9.2 degrees off-set w.r.t payload module axis,
 - FoV $\sim 2250 \text{ deg}^2$ ($\sim 47 \text{ degrees} \times 47 \text{ degrees}$), the central area being observed by all 24 cameras
 - Two fast cameras used as Fine Guidance Sensor (FGS) in AOCS (performances much better than standard star trackers)

Payload configuration

- ❑ Multi-telescope configuration with 24 'normal' cameras with CCD-based focal planes, in the visible wavelength range (500-1000 nm)
- ❑ Normal cameras monitor stars of magnitude $m_v > 8$ (cadence of 25 s)
- ❑ Two 'fast' cameras observe brighter stars ($m_v \sim 4-8$) at high cadence (2.5 s)

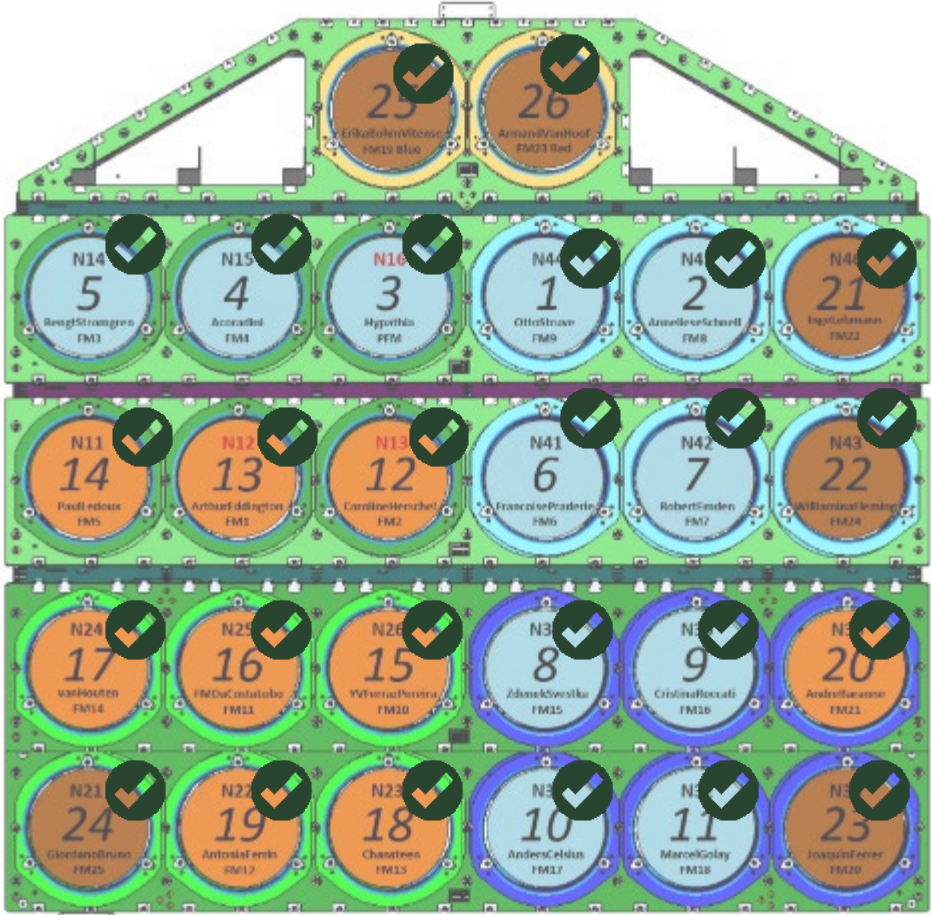


Architecture

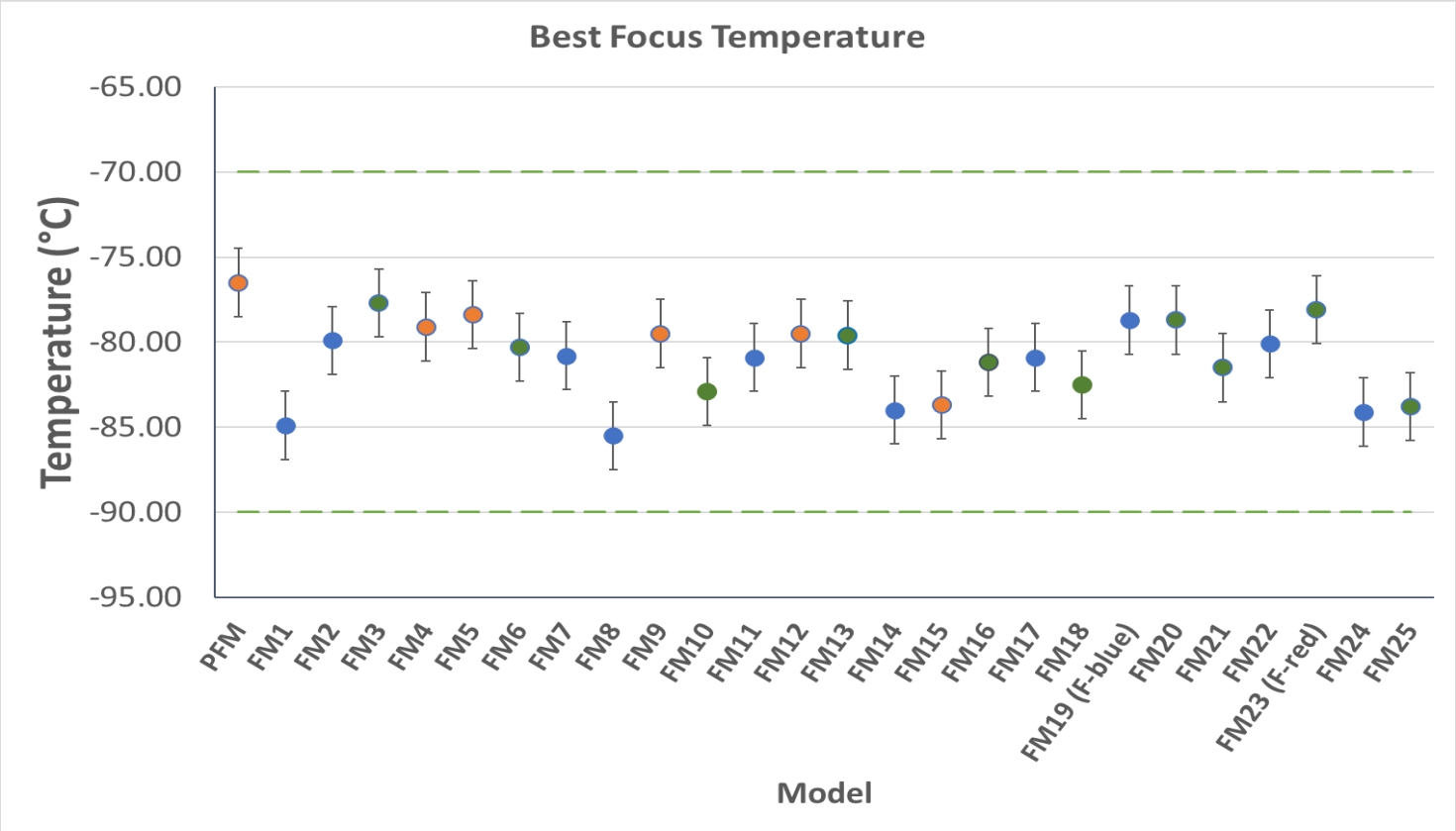


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PLATO Camera status



✓ All (24) N-CAMs and (2) F-CAMs tested, delivered to OHB prime and installed on OBA within pointing spec.

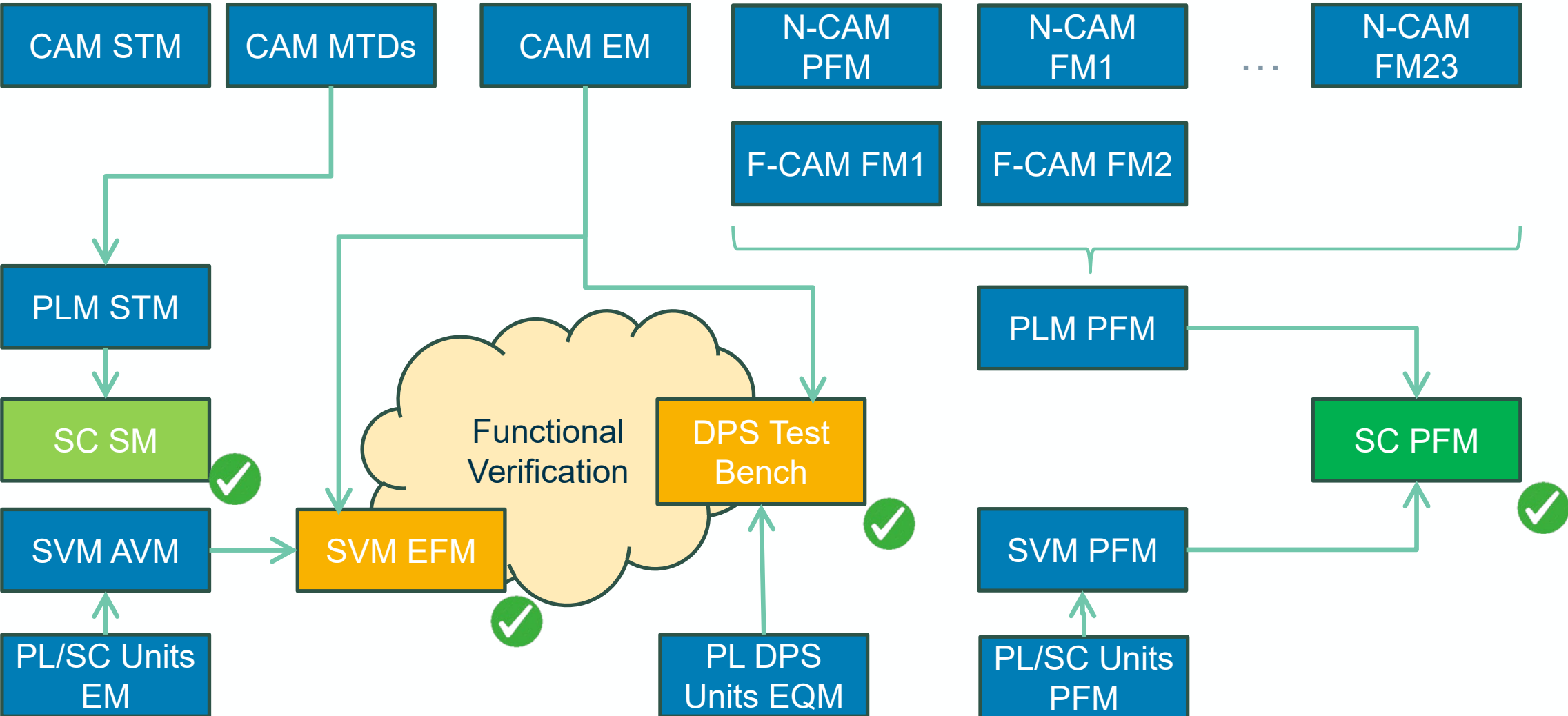




-
- Figure 10 consists of two panels. The left panel is a log-linear plot showing the Noise-to-Signal Ratio [ppm] on the y-axis (log scale from 10^0 to 10^4) versus the G0 Reference Star m_v on the x-axis (linear scale from 6 to 16). The plot includes curves for Instrument Noise (solid green), Photon Noise (dashed purple), Other Random Noise (dotted blue), Systematic Noise (dash-dot yellow), and Requirement (solid red). A vertical dashed line is at $m_v \approx 11.5$. The right panel is a heatmap showing the Noise to Signal Ratio [ppm] for $m_v = 11$ G0 Reference Star in 1h. The color scale ranges from 30 (blue) to 90 (red). The heatmap shows a central region with lower noise (blue) and outer regions with higher noise (red/orange). Numerical values are labeled in the heatmap cells: 43 in the center, 49 in the inner ring, 60 in the middle ring, and 84 in the outer ring.

Table 2-1: Mission performance requirements including compliance statements.

Payload Development Cycle



 **DPS System:**

- 
[→ THE EUROPEAN SPACE AGENCY](#)





End presentation

