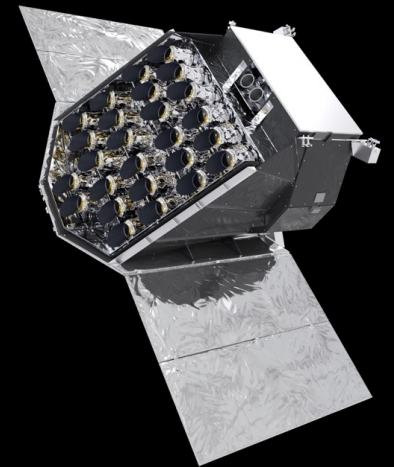


The ESPRESSO follow-up of small transiting exoplanets: perspectives for PLATO

F. Bouchy, M. Hobson, C. Lovis, F. Pepe, X. Dumusque, S. Udry, B. Lavie,
and the ESPRESSO consortium



ESPRESSO in a nutshell

Echelle Spectrograph for Rocky Exoplanets and Stable Spectroscopic Observation

collect the light of any UT
independently or together

Unit
Telescope
VLT

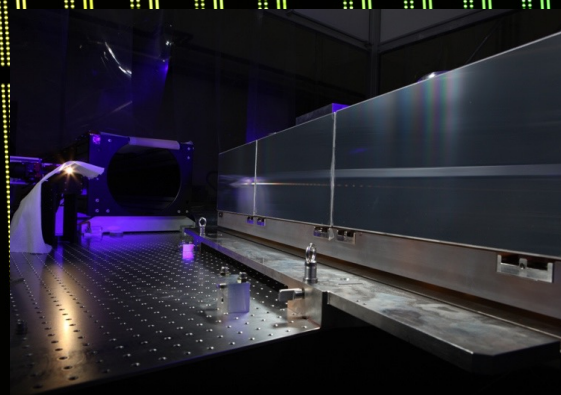
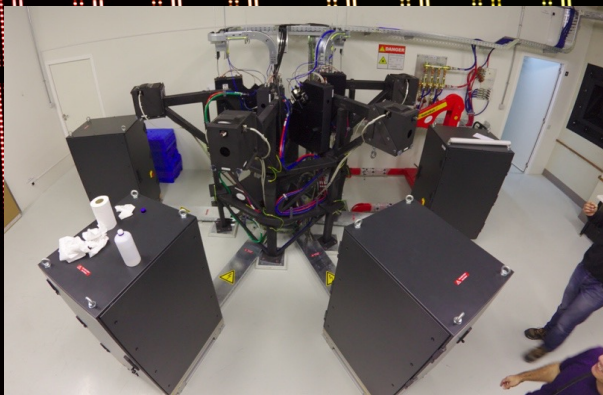
- 2007 STC and ESO identified need for a “HARPS” on the VLT
- Jan 2011 : Project Kick-Off
- Nov 2017 : First light at Paranal
- Oct 2018 : Start of scientific operations
- Apr 2023 : End of GTO
- second most demanded instrument on the VLT

Parameter	singleUHR	singleHR	multiMR
Wavelengths	Blue arm: 378 – 520 nm Red arm: 520 – 788 nm		
Resolving power	>190'000	138'000	70'000
Aperture on sky	0.5 arcsec	1.0 arcsec	4x1.0 arcsec
Spectral sampling	2.5 pixels	4.5 pixels	10 pixels

RV photon-noise uncertainty for Texp=15min

Spec-Type	V=11
G2V	98 cm/s
G8V	74 cm/s
K5V	62 cm/s
M2V	54 cm/s

Pepe et al. 2021

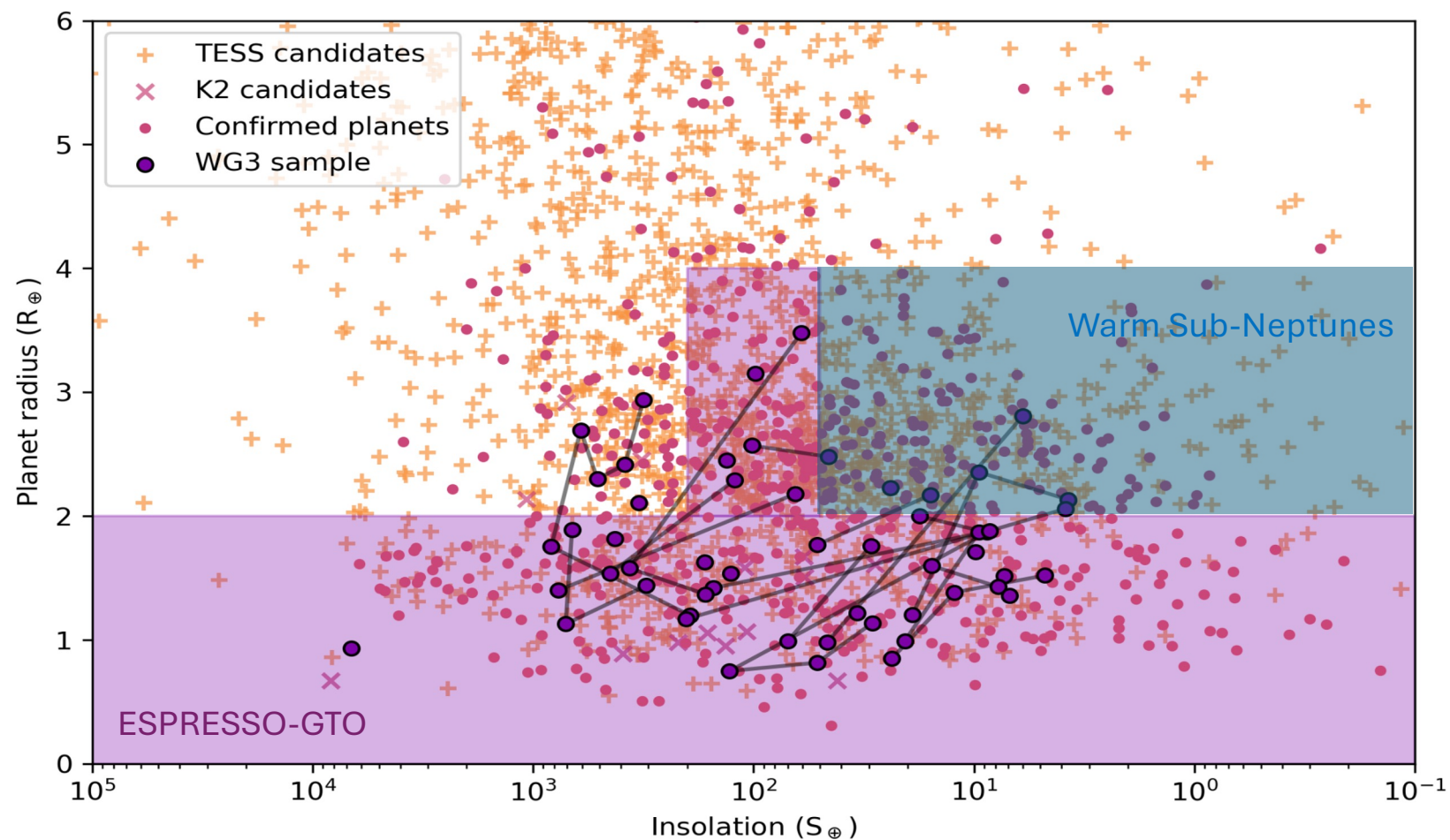




ESPRESSO-GTO WP3 science goals (PI F. Pepe) ~ 12 nights / semester x 8 semesters

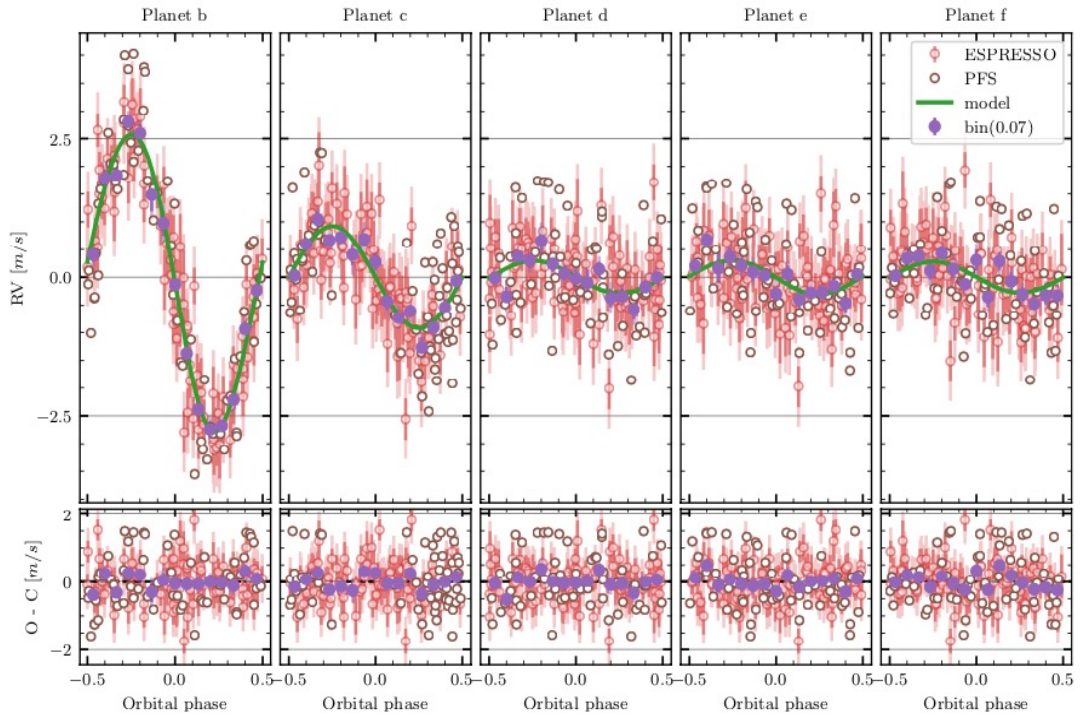
WP3.1) Characterize rocky population up to the habitable zone

WP3.2) Explore rocky/sub-Neptune transition within the 50-200x Earth irradiation



HD 23472: a multi-planetary system with three super-Earths and two potential super-Mercuries^{★,★★}

Barros et al. 2022



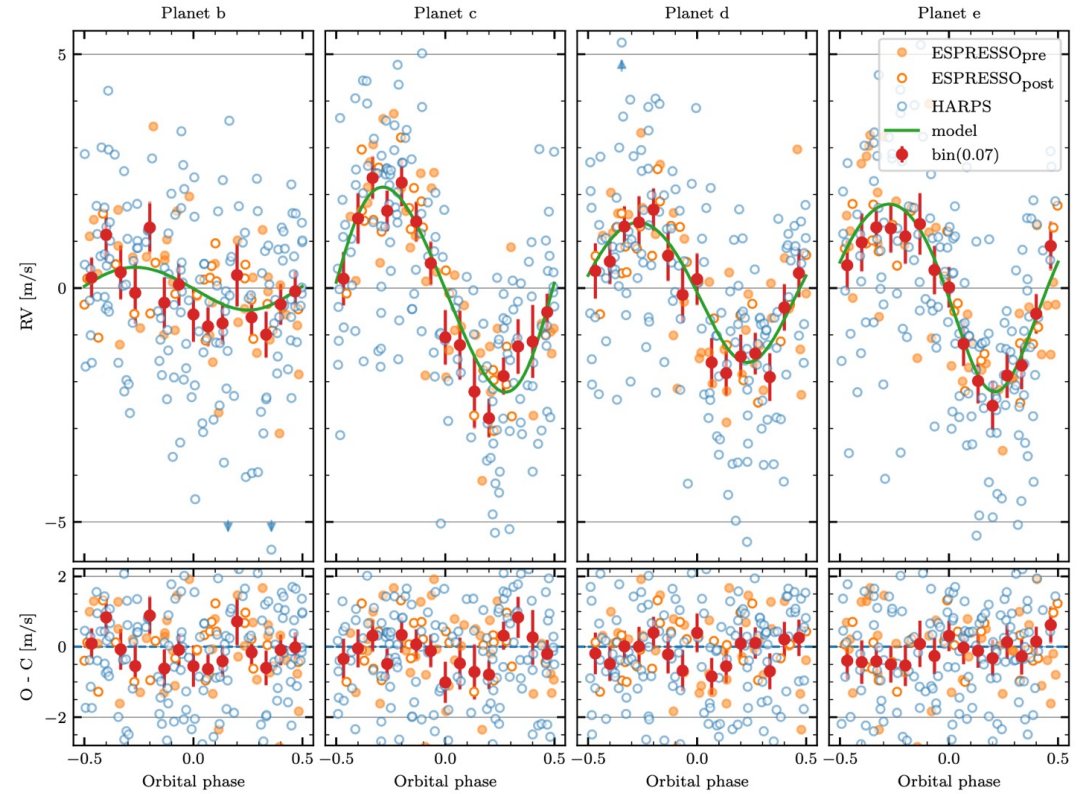
$8.3M_{\oplus}$ @ 17.7d | $3.4M_{\oplus}$ @ 29.8d | $0.55M_{\oplus}$ @ 4.0d | $0.72M_{\oplus}$ @ 7.9d | $0.77M_{\oplus}$ @ 12.2d

HD23472 K4V V=9.7

104 ESPRESSO + 64 PFS 2.3 years

Warm terrestrial planet with half the mass of Venus transiting a nearby star^{★,★★}

Demangeon et al. 2021

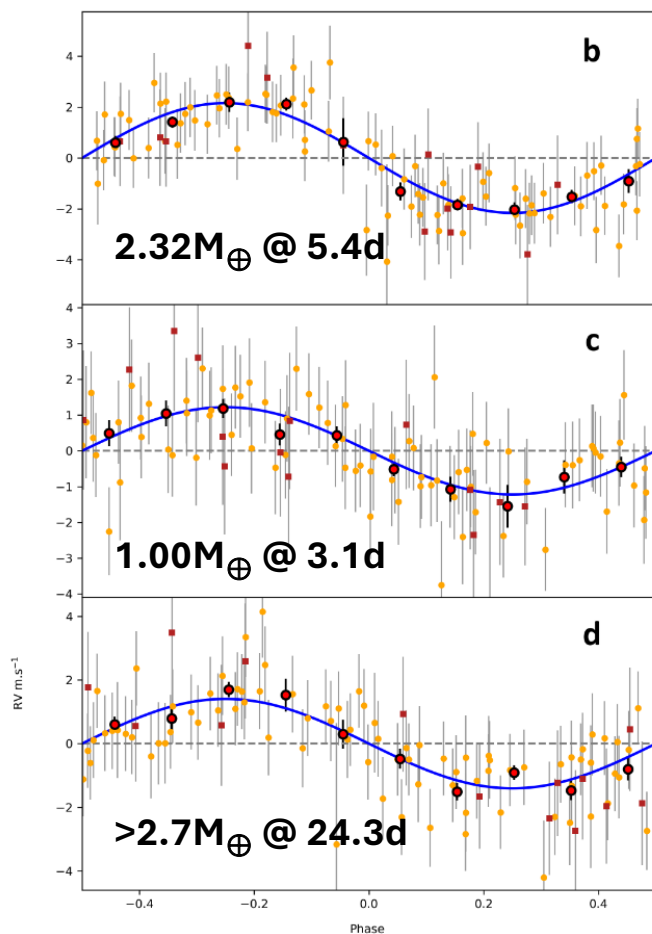


$0.40M_{\oplus}$ @ 2.25d | $2.22M_{\oplus}$ @ 3.69d | $1.94M_{\oplus}$ @ 7.45d | $>3.06M_{\oplus}$ @ 12.8d

L98-59 M3V V=11.7

66 ESPRESSO + 165 HARPS 1.5 years

Planetary system around LTT 1445A unveiled by ESPRESSO:
Multiple planets in a triple M-dwarf system^{★,★★}



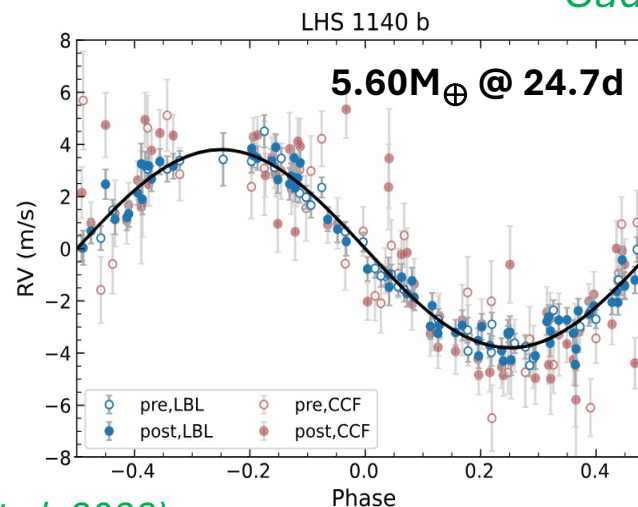
LTT1445A M3V V=10.5
84 ESPRESSO 1.7 years

Lavie et al. 2023

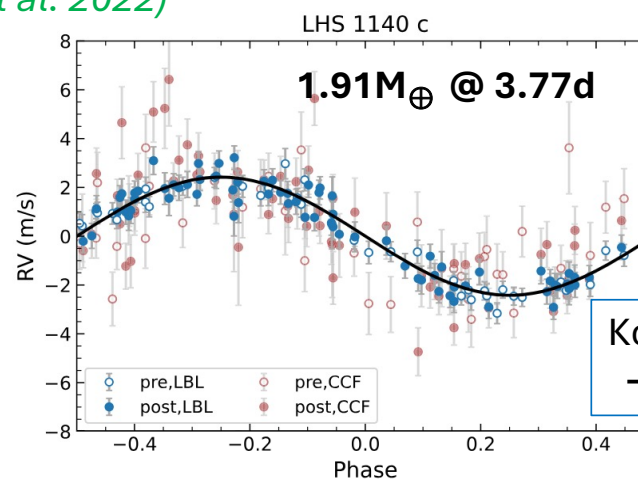
*see also poster from Y. Eschen about Gl12b
and talk from S. Grouffal about HIP41378*

New Mass and Radius Constraints on the LHS 1140 Planets

Cadieux et al. 2024



LBL (*Artigau et al. 2022*)



$K_c = 2.22 \pm 0.20 \text{ m.s}^{-1}$
 $\rightarrow 2.42 \pm 0.07 \text{ m.s}^{-1}$

LHS1140 M4.5 V=14.1
117 ESPRESSO 1.2 years

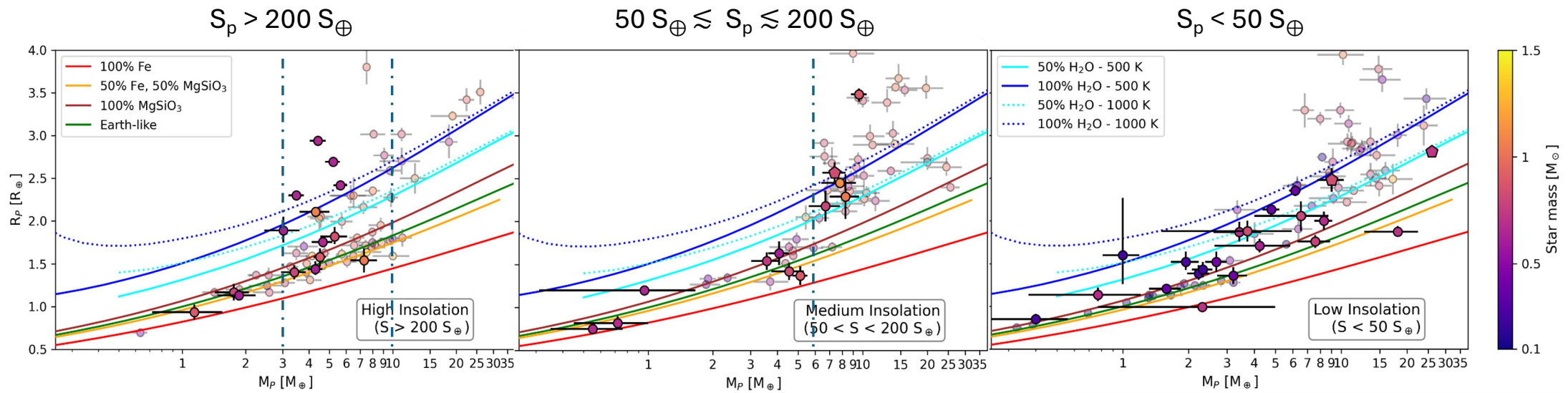
ESPRESSO Harvest

Hobson et al. 2025, in prep

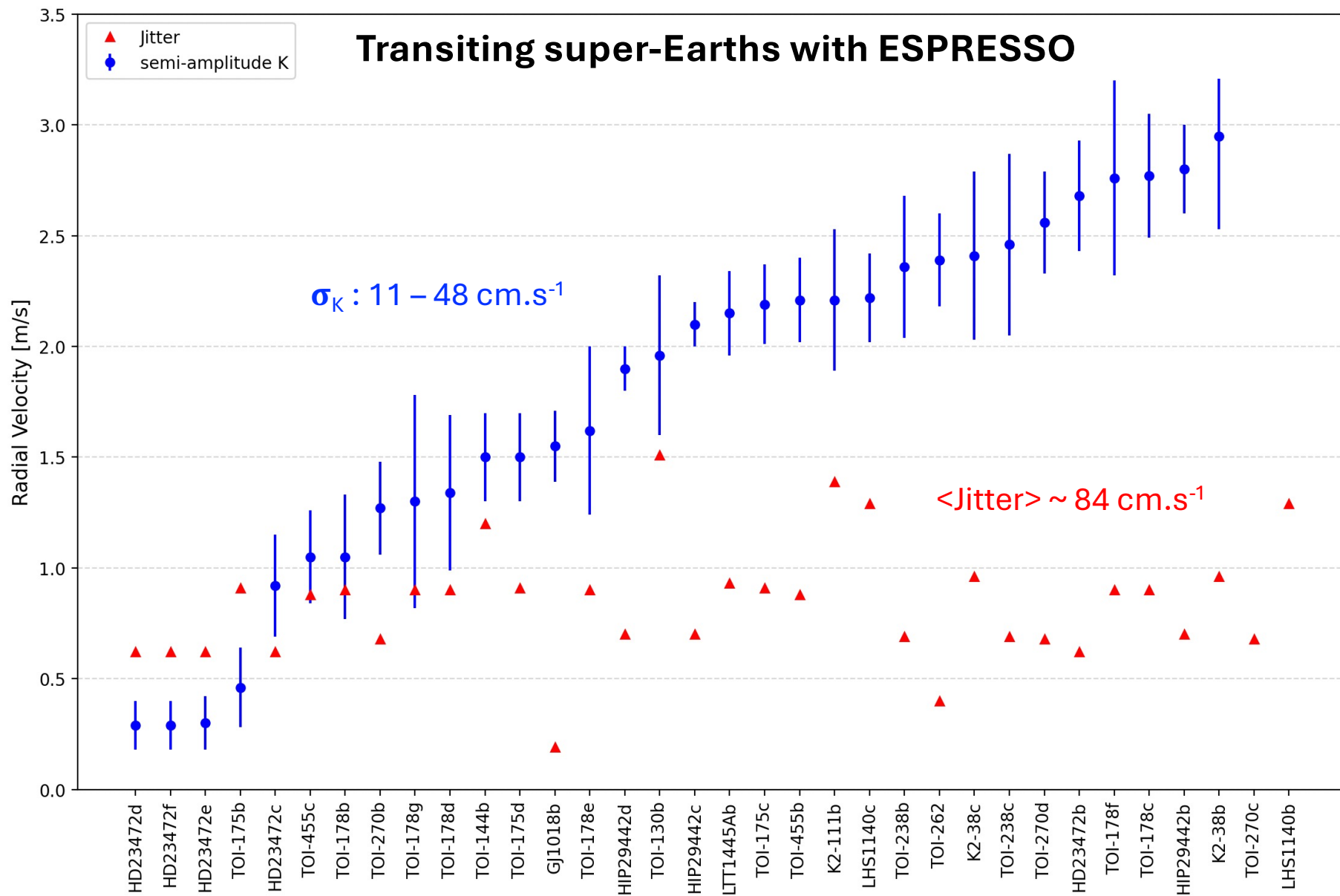
52 planets in 26 systems from ESPRESSO-GTO WP3

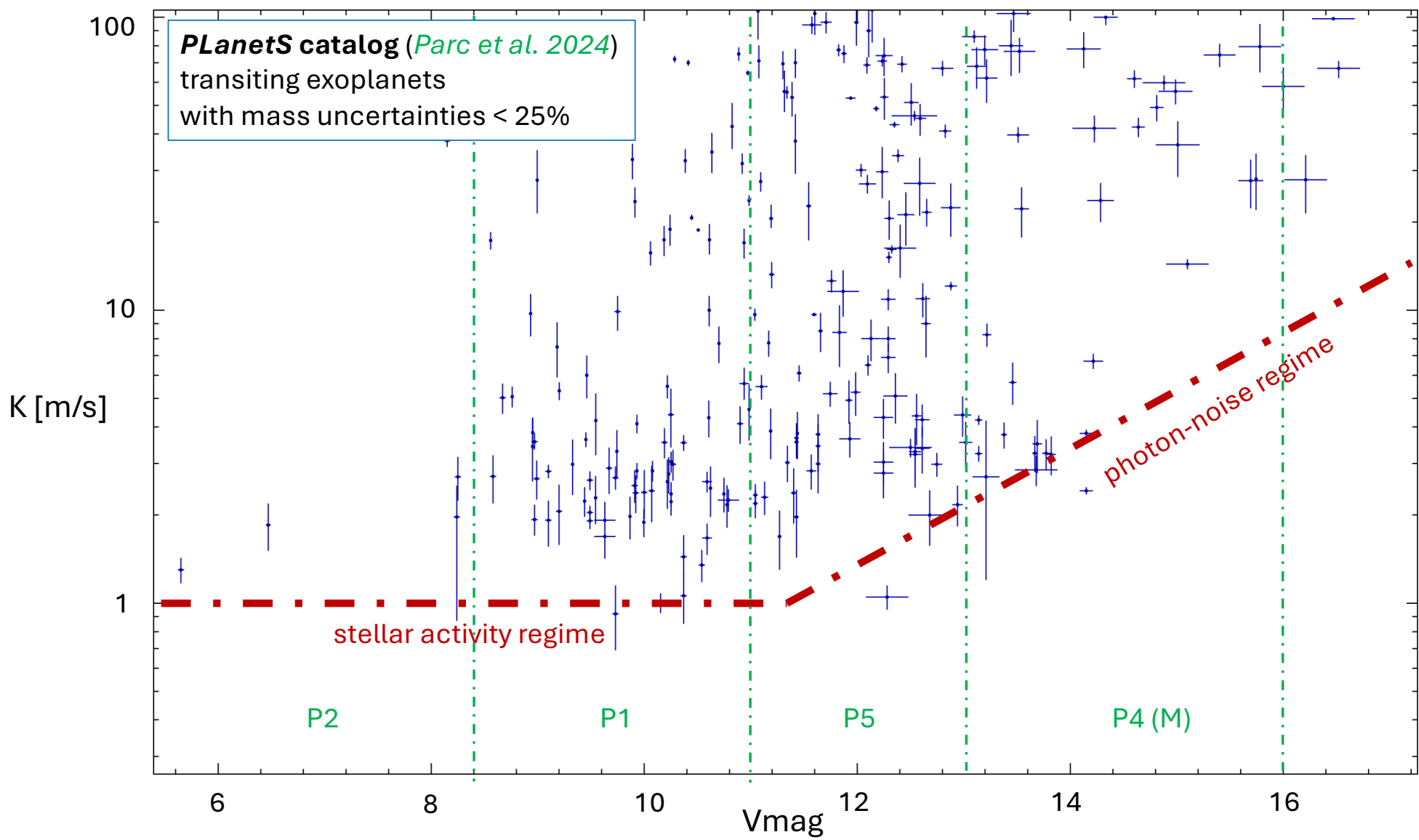
Studies of rocky-to-volatile transition for different regime of insolation

Comparison with the *PlanetS* catalog (*Parc et al. 2024*)



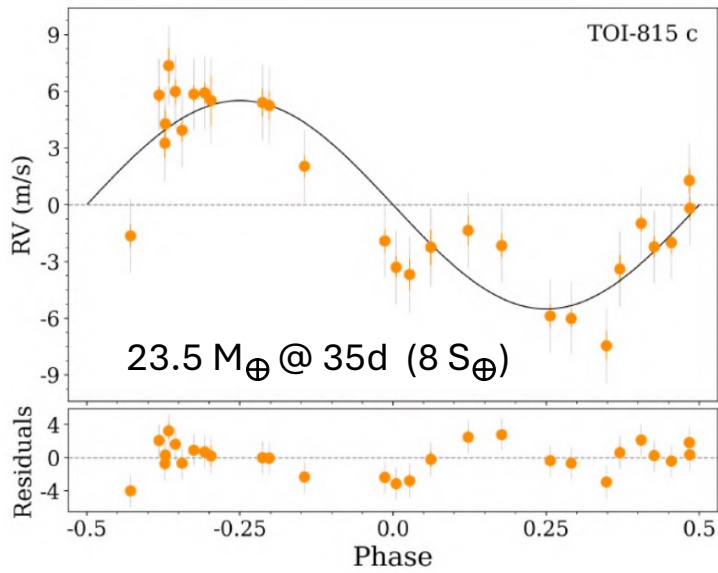
K





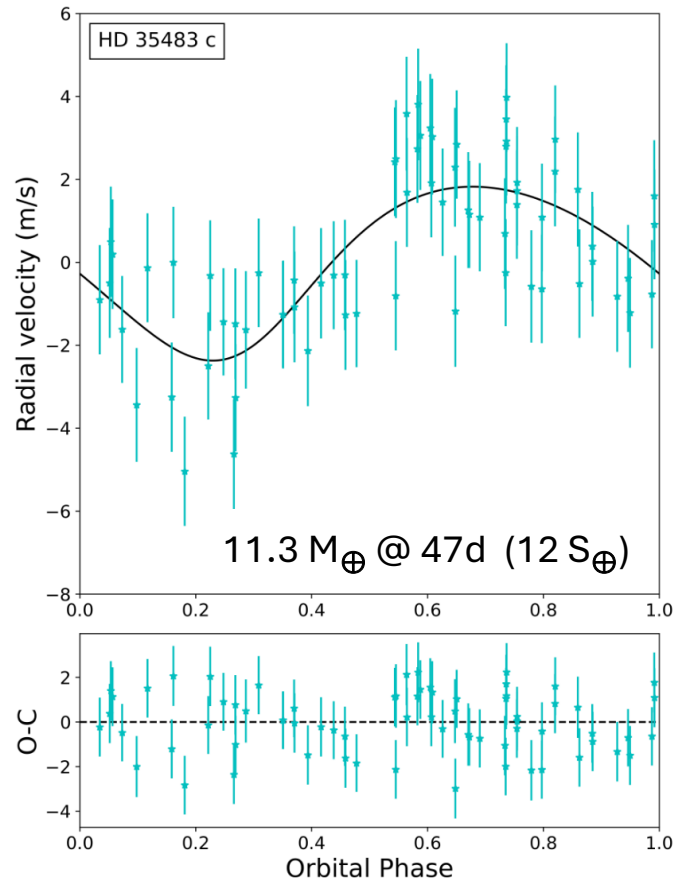
Warm sub-Neptunes with ESPRESSO

Psaridi et al. 2024



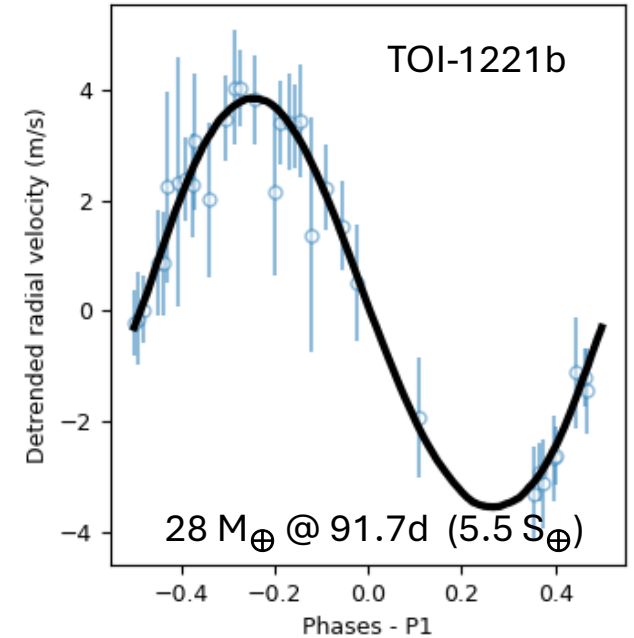
TOI-815 K3V V=10.2
34 ESPRESSO 1.7 years

Hesse et al. 2025 in press

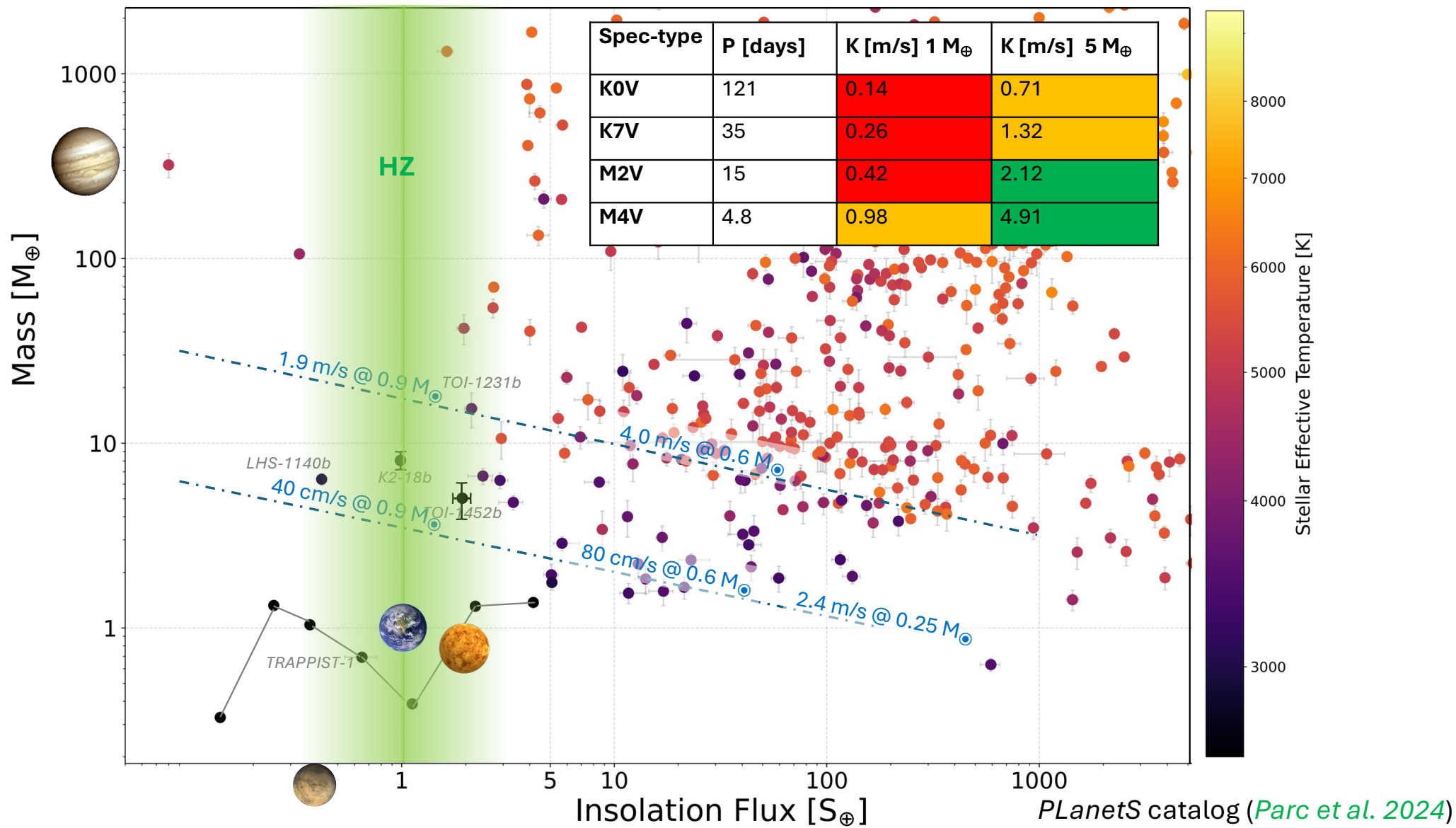


HD35483 G5V V=9.4
70 ESPRESSO 0.5 year

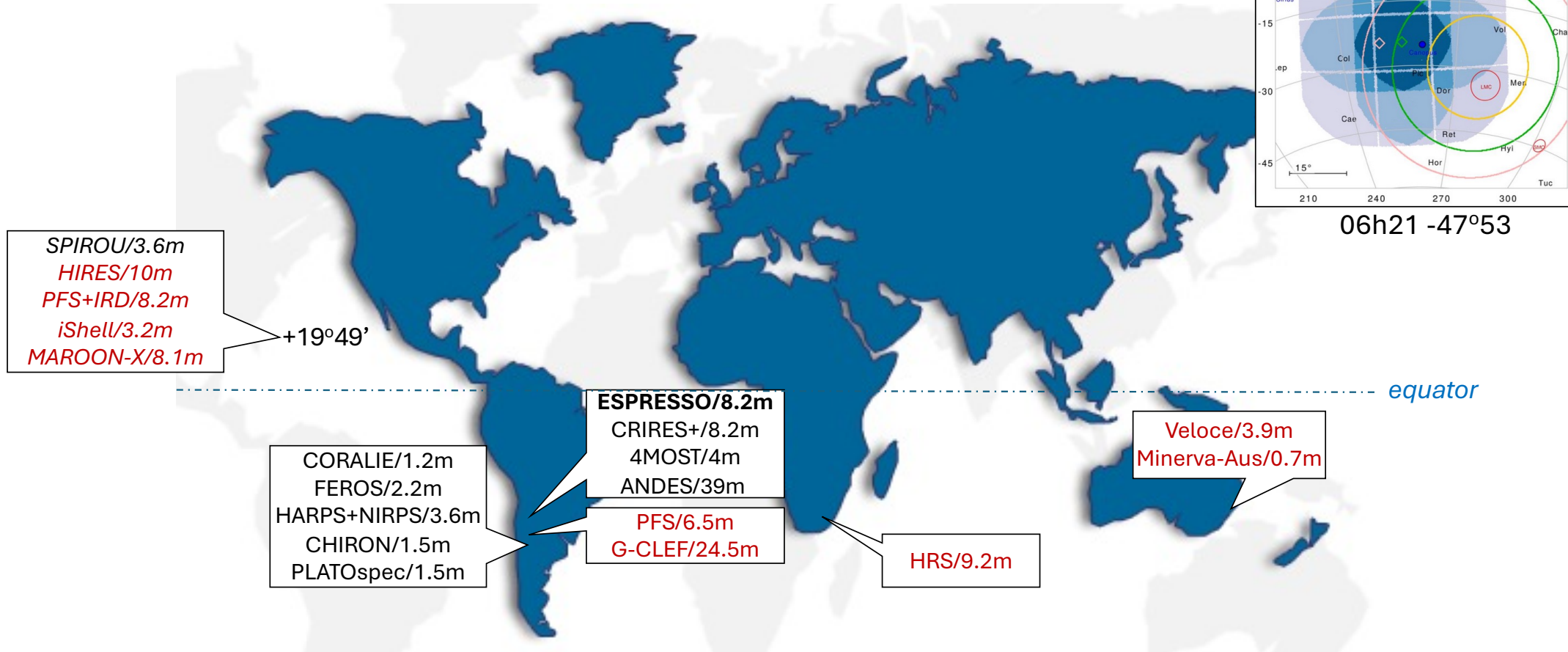
Bouchy et al. in prep



TOI-1221 G8V V=10.5
33 ESPRESSO 1.0 year



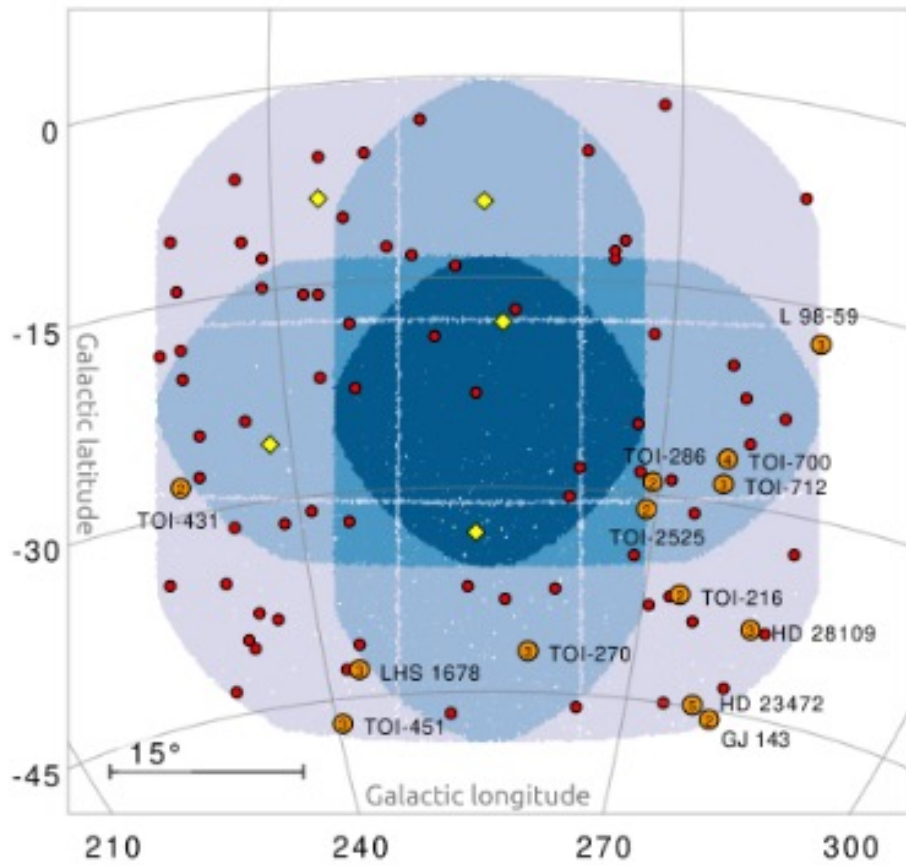
ESPRESSO will be key for PLATO LOPS2 follow-up



Non-European facilities

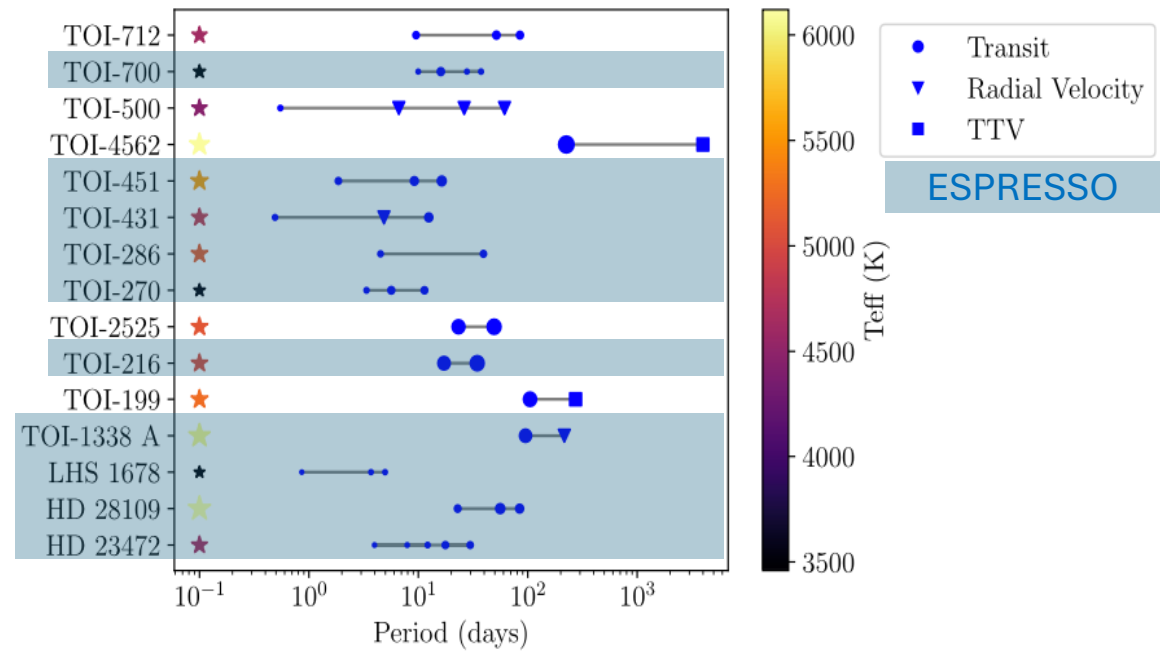
LOPS2 Center : Obs lat < +6° for airmass < 1.7

LOPS2 edge 24 cameras : Obs lat < +17° for airmass < 1.7



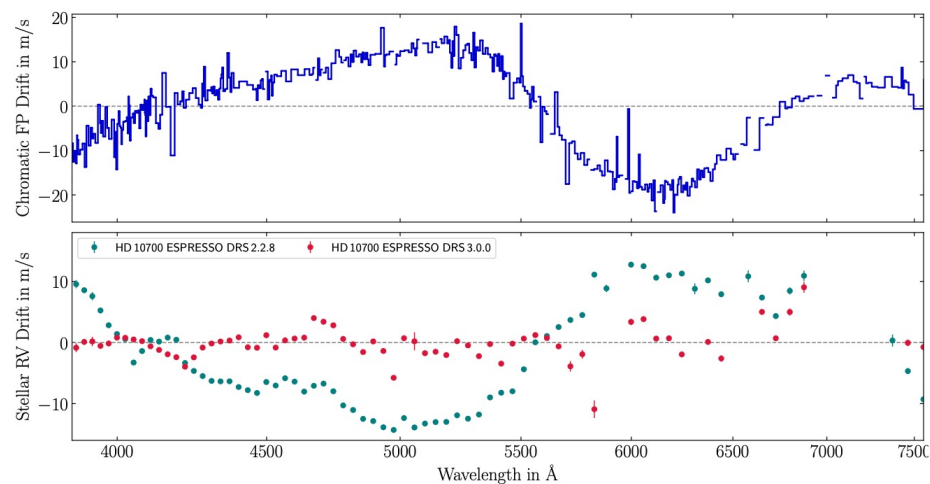
108 confirmed transiting exoplanets

Nascimbeni et al. 2025



15 systems with multiple exoplanets
(10 of them monitored by ESPRESSO)

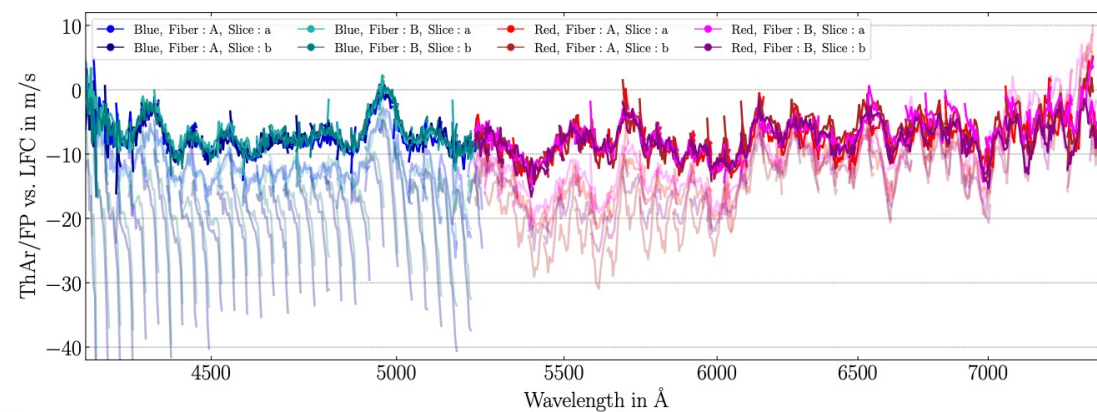
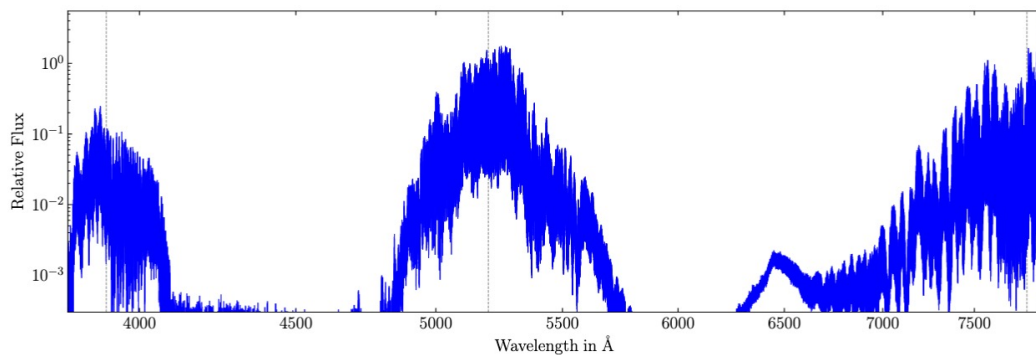
Eschen et al. 2024



**Continuous improvement of calibration sources
and Data Reduction Software**

Chromatic drift of the ESPRESSO Fabry-Pérot
Schmidt et al. 2022

Impact of modeling the instrument LSF
Schmidt & Bouchy 2024



Recent development BLUVES-LFC tested on ESPRESSO
Schmidt et al. in prep



ESPRESSO will be key for the PLATO follow-up of small transiting exoplanets

On-sky RV precision of better 40 cm.s^{-1} over 3.5 yr (*Figueira et al. 2025 in press*)

A lot of ESPRESSO archival data on multi-planetary systems part of LOPS2

PLATO+ESPRESSO : a niche for temperate exoplanets around K dwarfs

Stellar activity mitigation is fundamental

(*talk L. Malavolta + talk G.J. Talens + talk S. Aigrain & N. Meunier*)

Combination of TTVs + ESPRESSO RVs to derive masses of systems close to MMR

(*talk R. Mardling + talk A. Leleu + talk L. Pietro*)

On-going developement/upgrade on LFC (for long term accuracy)

Continuous upgrade of data reduction software and RV post-processing

(*talk of A. Silva , talk of M. Crétignier & N. Hara*)